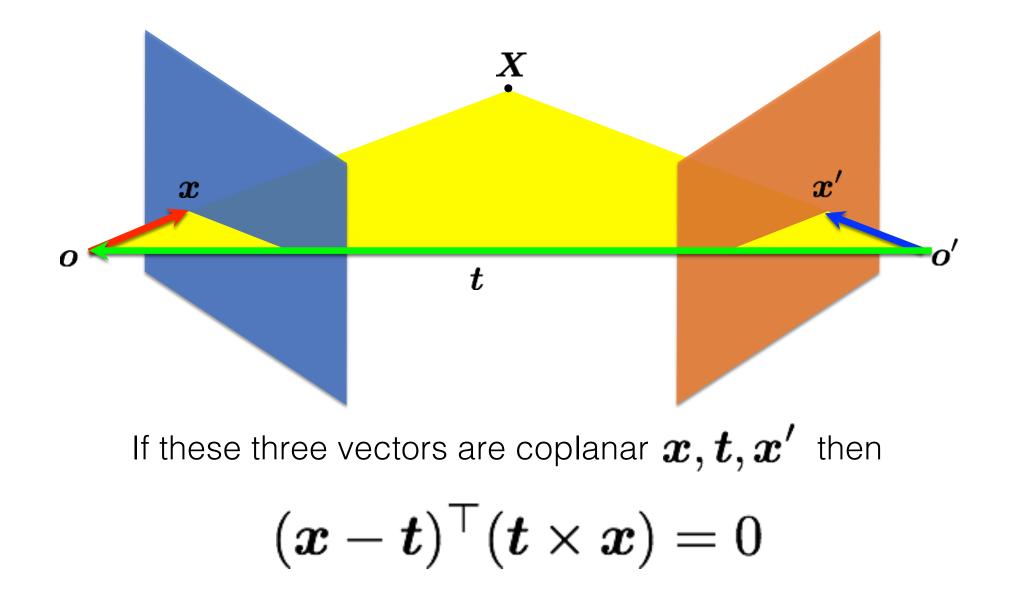
#### Stereo II



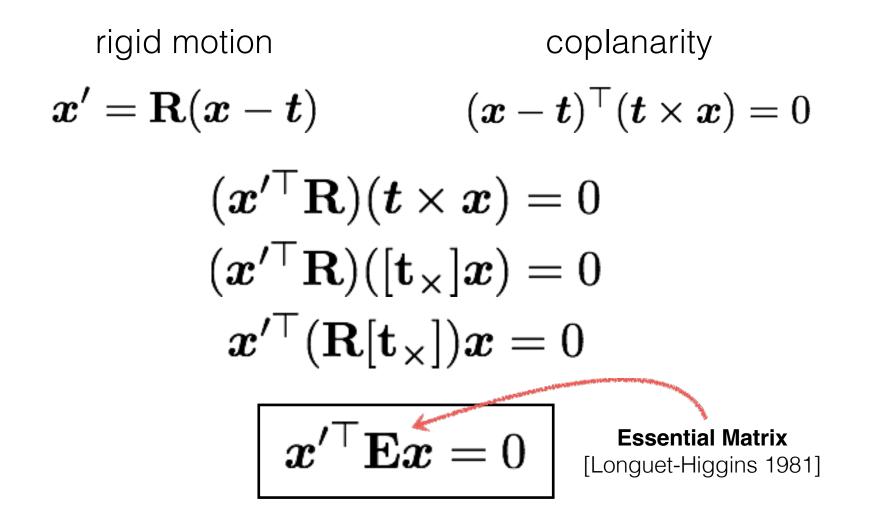
CSC420 David Lindell University of Toronto <u>cs.toronto.edu/~lindell/teaching/420</u> Slide credit: Babak Taati ←Ahmed Ashraf ←Sanja Fidler, Yannis Gkioulekas, Kris Kitani, Srinivasa Narasimhan







#### **Essential Matrix**



 $\hat{\boldsymbol{x}}^{\prime \top} \mathbf{E} \hat{\boldsymbol{x}} = 0$ 

The essential matrix operates on image points expressed in **2D coordinates expressed in the camera coordinate system** 

$$\hat{m{x}'}=\!\mathbf{K}'^{-1}m{x}'$$



point

image point

Writing out the epipolar constraint in terms of image coordinates

$$\mathbf{K}^{\prime - \top} \mathbf{E} \mathbf{K}^{-1} \mathbf{x} = 0$$
$$\mathbf{x}^{\prime \top} (\mathbf{K}^{\prime - \top} \mathbf{E} \mathbf{K}^{-1}) \mathbf{x} = 0$$
$$\mathbf{x}^{\prime \top} \mathbf{F} \mathbf{x} = \mathbf{0}$$

Longuet-Higgins equation

$$oldsymbol{x}'^{ op} oldsymbol{E} oldsymbol{x} = 0$$

Epipolar lines 
$$egin{array}{ccc} m{x}^{ op}m{l}=0 & m{x}'^{ op}m{l}'=0 \ m{E}m{x} & m{l}=m{E}^Tm{x}' \end{array}$$

Epipoles 
$$e'^{ op} \mathbf{ar{E}} = \mathbf{0}$$
  $\mathbf{ar{E}} e = \mathbf{0}$ 

(points in **image** coordinates)

Breaking down the fundamental matrix

# $\mathbf{F} = \mathbf{K}'^{-\top} \mathbf{E} \mathbf{K}^{-1}$ $\mathbf{F} = \mathbf{K}'^{-\top} [\mathbf{t}_{\times}] \mathbf{R} \mathbf{K}^{-1}$

Depends on both intrinsic and extrinsic parameters

Breaking down the fundamental matrix

# $\mathbf{F} = \mathbf{K}'^{-\top} \mathbf{E} \mathbf{K}^{-1}$ $\mathbf{F} = \mathbf{K}'^{-\top} [\mathbf{t}_{\times}] \mathbf{R} \mathbf{K}^{-1}$

Depends on both intrinsic and extrinsic parameters

How would you solve for F?

$$oldsymbol{x}_m^{\prime op} \mathbf{F} oldsymbol{x}_m = 0$$

### The 8-point algorithm

Assume you have *M* matched *image* points

$$\{\boldsymbol{x}_m, \boldsymbol{x}_m'\}$$
  $m = 1, \dots, M$ 

Each correspondence should satisfy

$$oldsymbol{x}_m^{\prime op} \mathbf{F} oldsymbol{x}_m = 0$$

How would you solve for the 3 x 3 **F** matrix?

Assume you have *M* matched *image* points

$$\{\boldsymbol{x}_m, \boldsymbol{x}_m'\}$$
  $m = 1, \dots, M$ 

Each correspondence should satisfy

$$oldsymbol{x}_m^{\prime op} \mathbf{F} oldsymbol{x}_m = 0$$

How would you solve for the 3 x 3 **F** matrix?

S V D

Assume you have *M* matched *image* points

$$\{\boldsymbol{x}_m, \boldsymbol{x}_m'\}$$
  $m = 1, \dots, M$ 

Each correspondence should satisfy

 $\boldsymbol{x}_m^{\prime op} \mathbf{F} \boldsymbol{x}_m = 0$ 

How would you solve for the 3 x 3 **F** matrix?

Set up a homogeneous linear system with 9 unknowns

$$oldsymbol{x}_m^{\prime \mid} \mathbf{F} oldsymbol{x}_m = 0$$
  
 $\left[ egin{array}{cccc} x_m^{\prime \mid} & y_m^{\prime \mid} & 1 \end{array} 
ight] \left[ egin{array}{cccc} f_1 & f_2 & f_3 \ f_4 & f_5 & f_6 \ f_7 & f_8 & f_9 \end{array} 
ight] \left[ egin{array}{cccc} x_m \ y_m \ 1 \end{array} 
ight] = 0$ 

#### How many equation do you get from one correspondence?

$$\begin{bmatrix} x'_{m} & y'_{m} & 1 \end{bmatrix} \begin{bmatrix} f_{1} & f_{2} & f_{3} \\ f_{4} & f_{5} & f_{6} \\ f_{7} & f_{8} & f_{9} \end{bmatrix} \begin{bmatrix} x_{m} \\ y_{m} \\ 1 \end{bmatrix} = 0$$

ONE correspondence gives you ONE equation

$$\begin{aligned} x_m x'_m f_1 + x_m y'_m f_2 + x_m f_3 + \\ y_m x'_m f_4 + y_m y'_m f_5 + y_m f_6 + \\ x'_m f_7 + y'_m f_8 + f_9 &= 0 \end{aligned}$$

$$\begin{bmatrix} x'_{m} & y'_{m} & 1 \end{bmatrix} \begin{bmatrix} f_{1} & f_{2} & f_{3} \\ f_{4} & f_{5} & f_{6} \\ f_{7} & f_{8} & f_{9} \end{bmatrix} \begin{bmatrix} x_{m} \\ y_{m} \\ 1 \end{bmatrix} = 0$$

Set up a homogeneous linear system with 9 unknowns

How many equations do you need?

Each point pair (according to epipolar constraint) contributes only one <u>scalar</u> equation

$$\boldsymbol{x}_m^{\prime op} \mathbf{F} \boldsymbol{x}_m = 0$$

**Note:** This is different from the Homography estimation where each point pair contributes 2 equations.

We need at least 8 points

## Hence, the 8 point algorithm!

How do you solve a homogeneous linear system?

## $\mathbf{A} \mathbf{X} = \mathbf{0}$

How do you solve a homogeneous linear system?

## $\mathbf{A} \mathbf{X} = \mathbf{0}$

Total Least Squares minimize  $\|\mathbf{A}\mathbf{x}\|^2$ 

subject to  $\| \boldsymbol{x} \|^2 = 1$ 

How do you solve a homogeneous linear system?

## $\mathbf{A} \mathbf{X} = \mathbf{0}$

Total Least Squares minimize  $\|\mathbf{A}\mathbf{x}\|^2$ subject to  $\|\mathbf{x}\|^2 = 1$ SVDD!

0. (Normalize points)

- 1. Construct the M x 9 matrix **A**
- 2. Find the SVD of  $\boldsymbol{\mathsf{A}}$
- 3. Entries of  ${\bf F}$  are the elements of column of

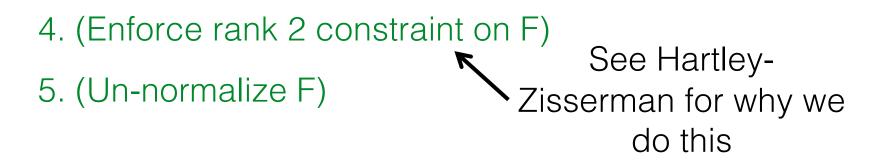
V corresponding to the least singular value

- 4. (Enforce rank 2 constraint on F)
- 5. (Un-normalize F)

0. (Normalize points)

- 1. Construct the M x 9 matrix **A**
- 2. Find the SVD of  $\boldsymbol{\mathsf{A}}$
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V corresponding to the least singular value

4. (Enforce rank 2 constraint on F)

5. (Un-normalize F)

How do we do this?

0. (Normalize points)

- 1. Construct the M x 9 matrix **A**
- 2. Find the SVD of  $\boldsymbol{\mathsf{A}}$
- 3. Entries of  ${\bf F}$  are the elements of column of

V corresponding to the least singular value

4. (Enforce rank 2 constraint on F)

5. (Un-normalize F)

How do we do this?

SVD!

## Enforcing rank constraints

Problem: Given a matrix F, find the matrix F' of rank k that is closest to F,

$$\min_{F'} ||F - F'||^2$$
$$\operatorname{rank}(F') = k$$

Solution: Compute the singular value decomposition of F,

$$F = U\Sigma V^T$$

Form a matrix  $\Sigma$ ' by replacing all but the k largest singular values in  $\Sigma$  with 0.

Then the problem solution is the matrix **F'** formed as,

$$F' = U\Sigma' V^T$$

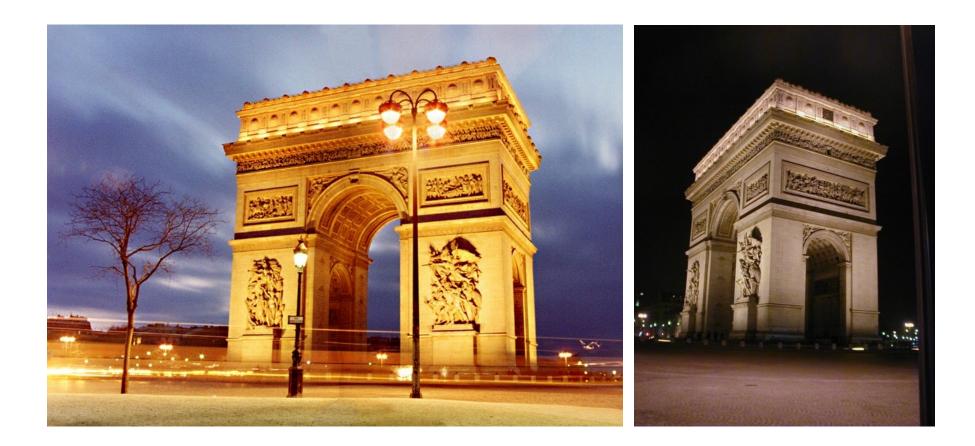
0. (Normalize points)

- 1. Construct the M x 9 matrix **A**
- 2. Find the SVD of  $\boldsymbol{\mathsf{A}}$
- 3. Entries of  ${\bf F}$  are the elements of column of

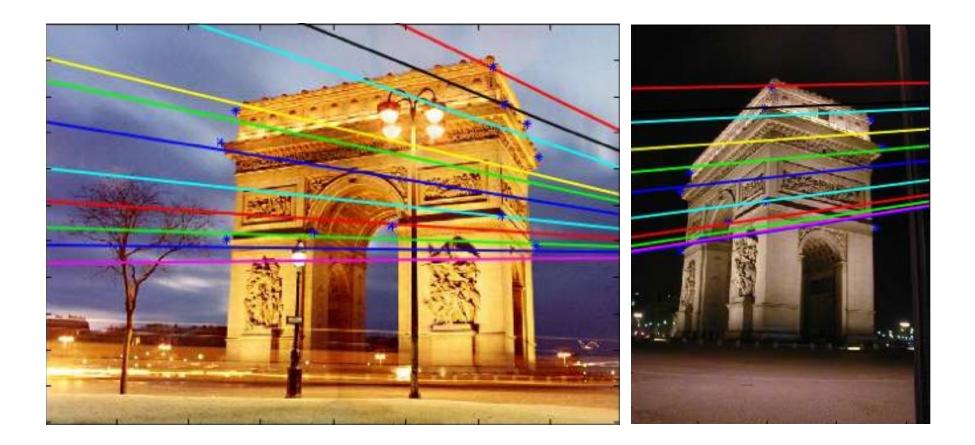
V corresponding to the least singular value

- 4. (Enforce rank 2 constraint on F)
- 5. (Un-normalize F)

## Example



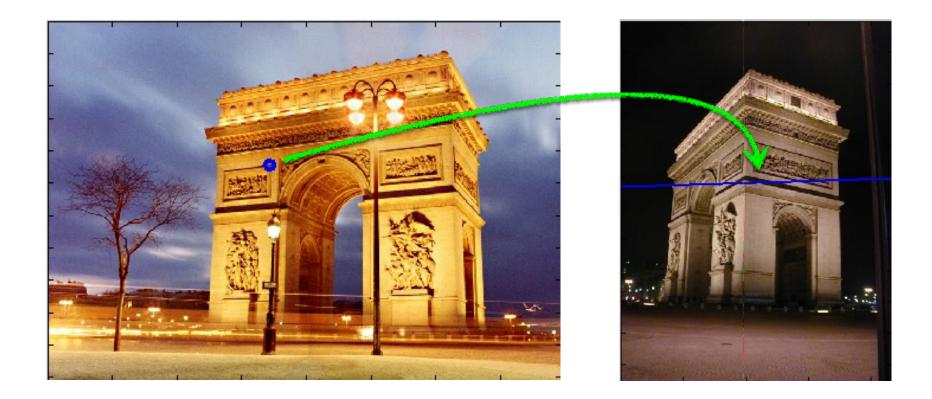
## epipolar lines



$$\mathbf{F} = \begin{bmatrix} -0.00310695 & -0.0025646 & 2.96584 \\ -0.028094 & -0.00771621 & 56.3813 \\ 13.1905 & -29.2007 & -9999.79 \end{bmatrix}$$

$$\mathbf{x} = \begin{bmatrix} 343.53\\ 221.70\\ 1.0 \end{bmatrix}$$
$$\mathbf{l}' = \mathbf{F}\mathbf{x}$$
$$= \begin{bmatrix} 0.0295\\ 0.9996\\ -265.1531 \end{bmatrix}$$

$${}^{\prime} = \mathbf{F} oldsymbol{x} \ = \left[ egin{array}{c} 0.0295 \ 0.9996 \ -265.1531 \end{array} 
ight]$$



### Where is the epipole?



How would you compute it?



## $\mathbf{F} \boldsymbol{e} = \boldsymbol{0}$

The epipole is in the right null space of  ${\bf F}$ 

How would you solve for the epipole?



## $\mathbf{F} \boldsymbol{e} = \boldsymbol{0}$

The epipole is in the right null space of  ${\bf F}$ 

How would you solve for the epipole?

## SVD!

## Revisiting triangulation

#### How would you reconstruct 3D points?



Left image



Right image

#### How would you reconstruct 3D points?



Left image



Right image

1. Select point in one image (how?)

### How would you reconstruct 3D points?



Left image



Right image

- 1. Select point in one image (how?)
- 2. Form epipolar line for that point in second image (how?)

## How would you reconstruct 3D points?



Left image

Right image

- 1. Select point in one image (how?)
- 2. Form epipolar line for that point in second image (how?)
- 3. Find matching point along line (how?)

## How would you reconstruct 3D points?



Left image

Right image

- 1. Select point in one image (how?)
- 2. Form epipolar line for that point in second image (how?)
- 3. Find matching point along line (how?)
- 4. Perform triangulation (how?)

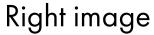
## Triangulation 3D point left image right image $\boldsymbol{x}$ x'() C'right camera with matrix $\mathbf{P'}$ left camera with matrix $\, {f P} \,$

## How would you reconstruct 3D points?



Left image





- 1. Select point in one image (how?)
- 2. Form epipolar line for that point in second image (how?)
- 3. Find matching point along line (how?)
- 4. Perform triangulation (how?)

What are the disadvantages of this procedure?

## Stereo rectification





#### What's different between these two images?







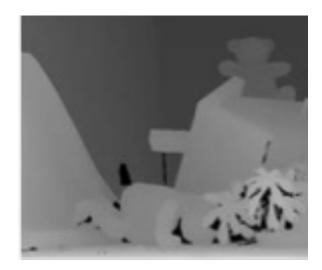


#### Objects that are close move more or less?

The amount of horizontal movement is inversely proportional to ...





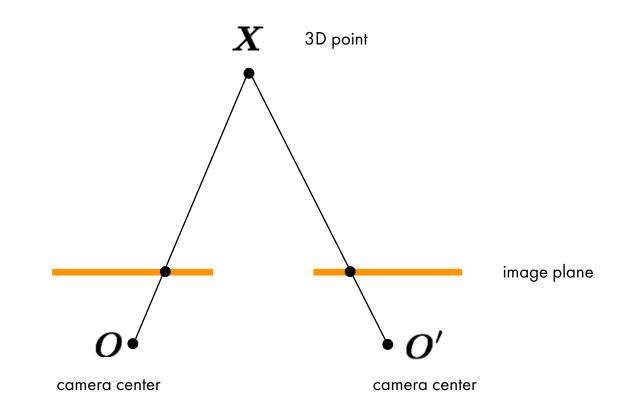


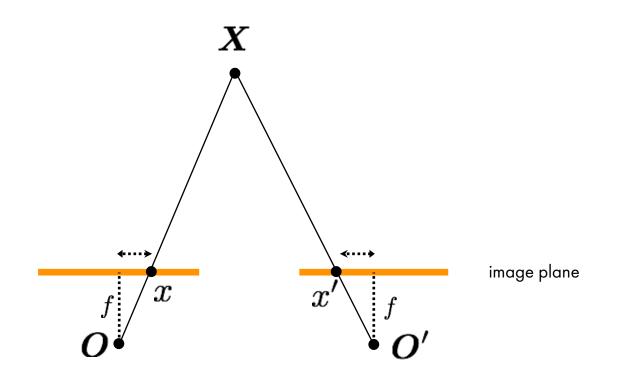
The amount of horizontal movement is inversely proportional to ...

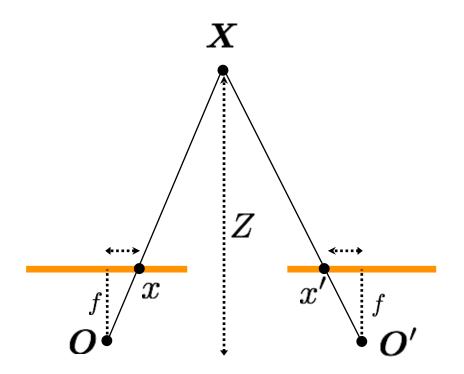


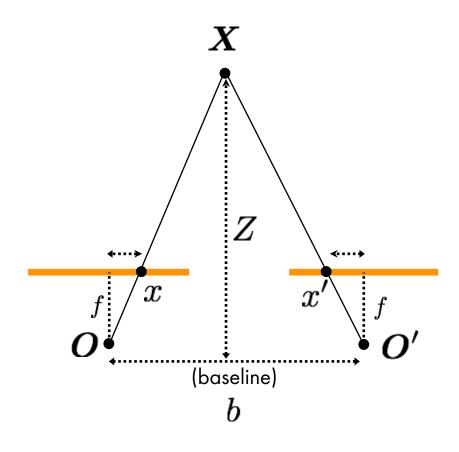
... the distance from the camera.

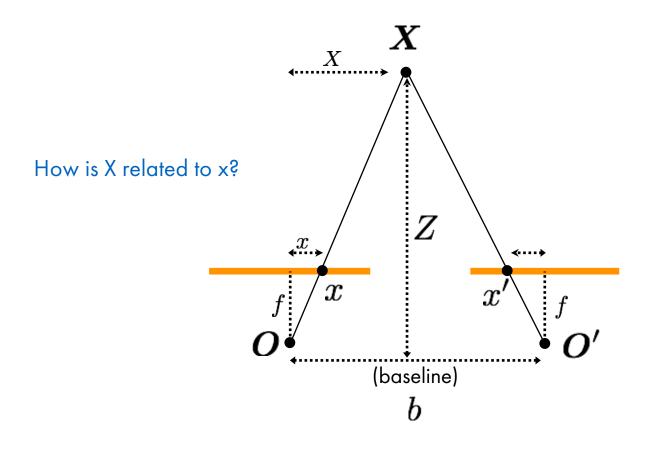


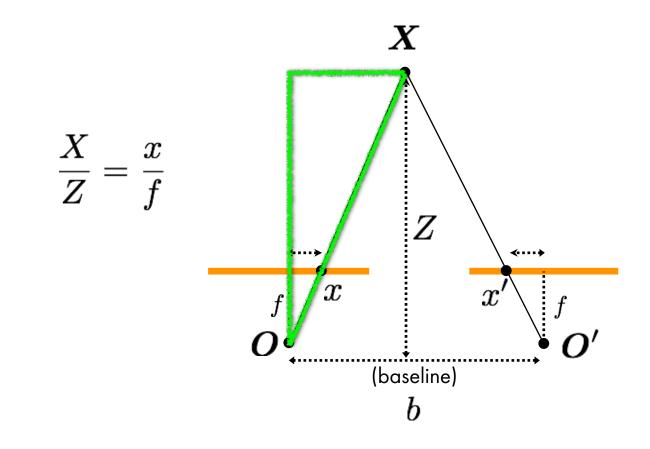


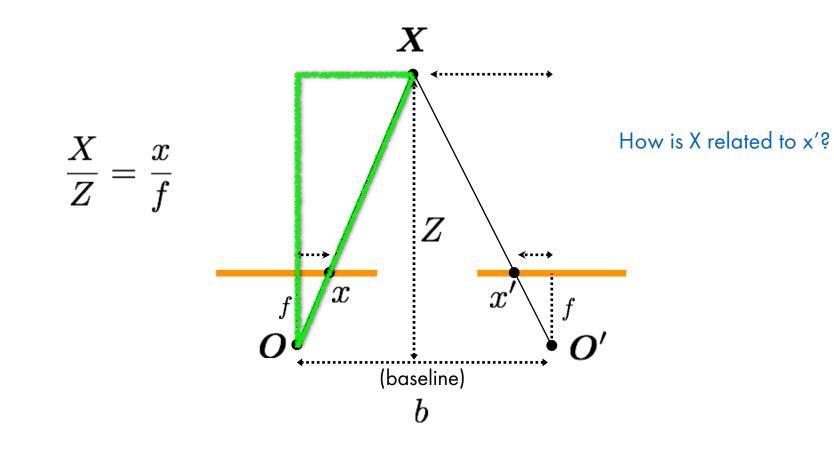


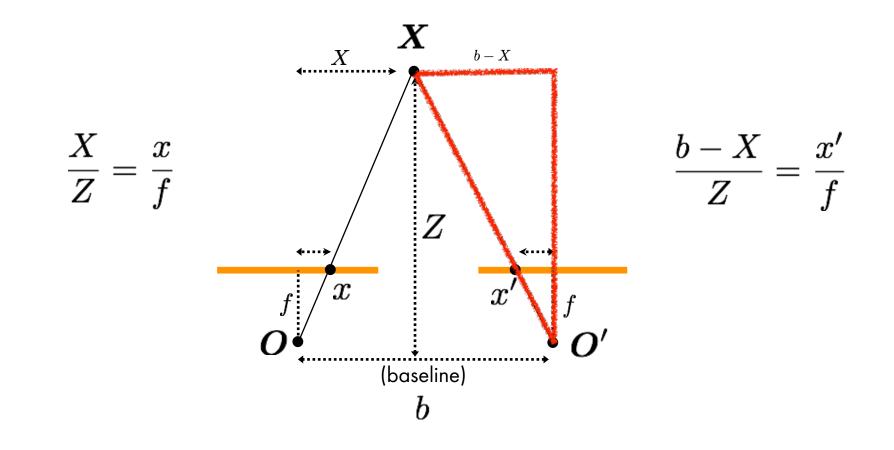


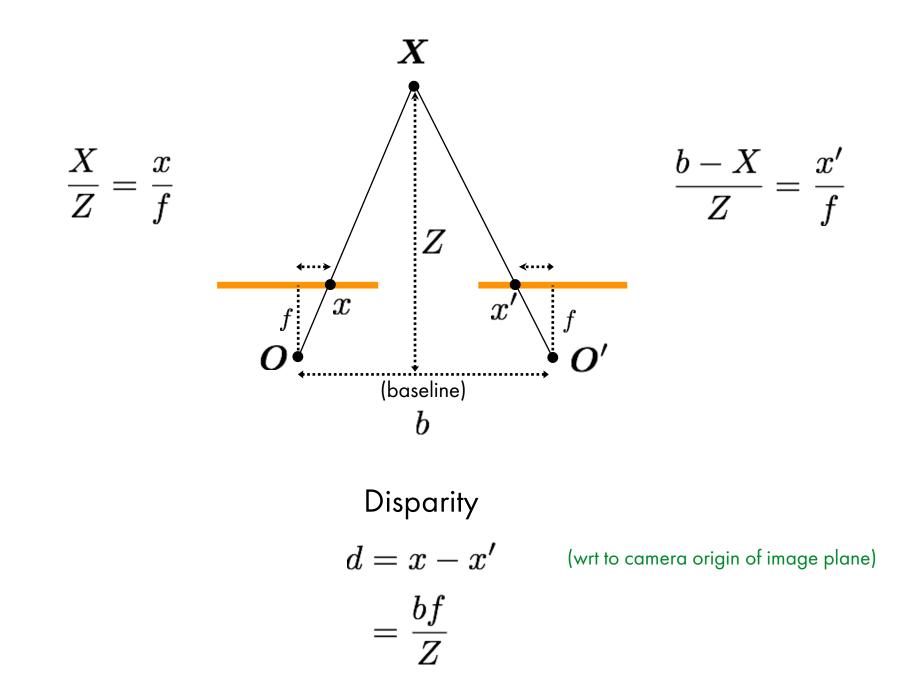


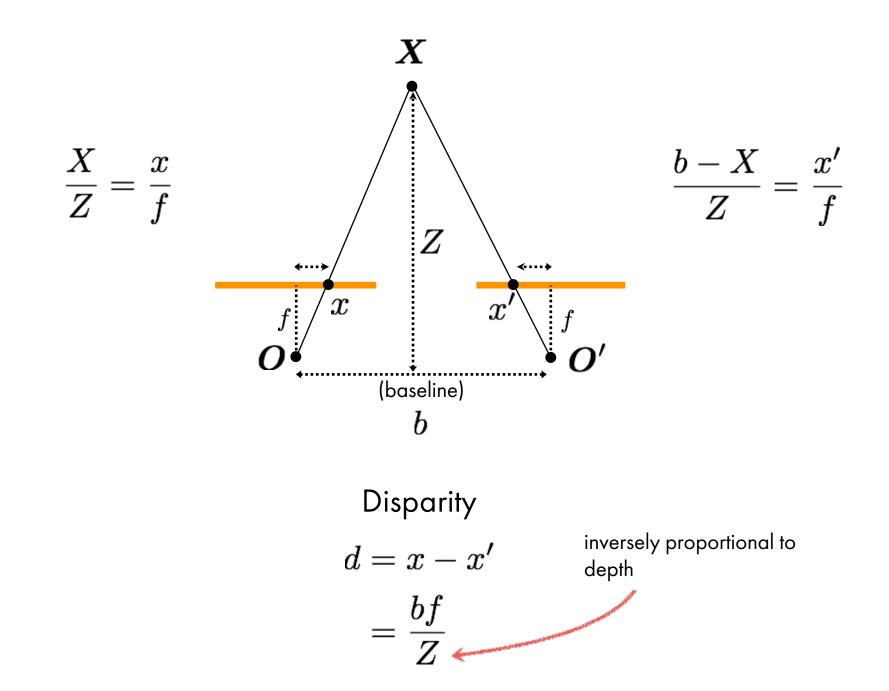














Subaru Eyesight system

#### Pre-collision braking

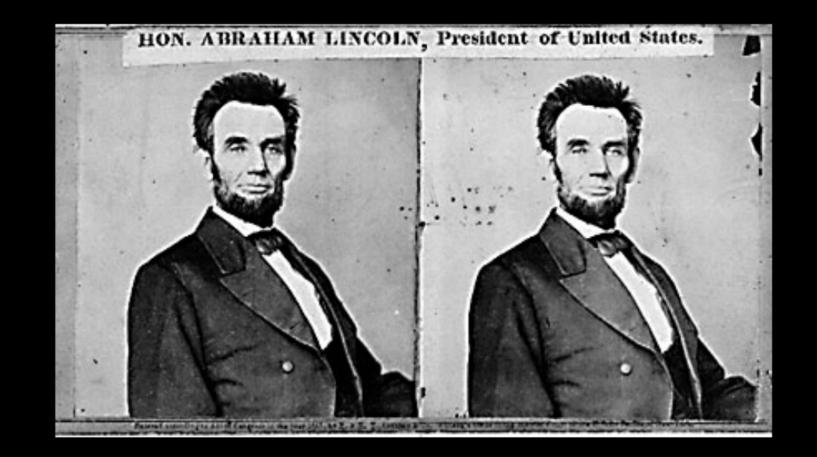


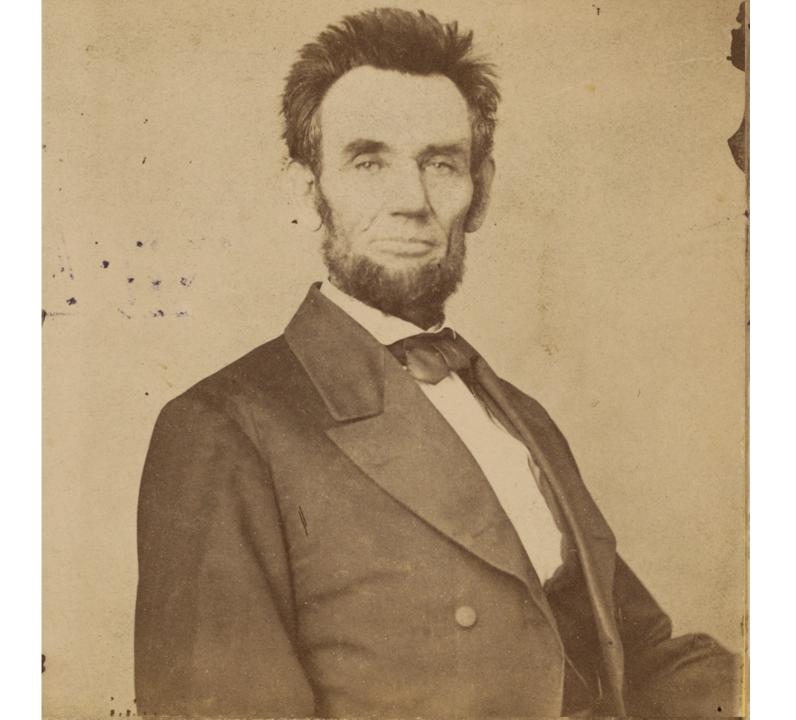


What other vision system uses disparity for depth sensing?

### Stereoscopes: A 19<sup>th</sup> Century Pastime









Public Library, Stereoscopic Looking Room, Chicago, by Phillips, 1923





Teesta suspension bridge-Darjeeling, India



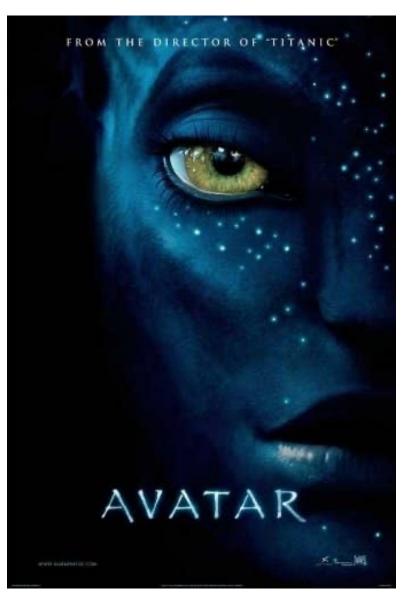


Mark Twain at Pool Table", no date, UCR Museum of Photography

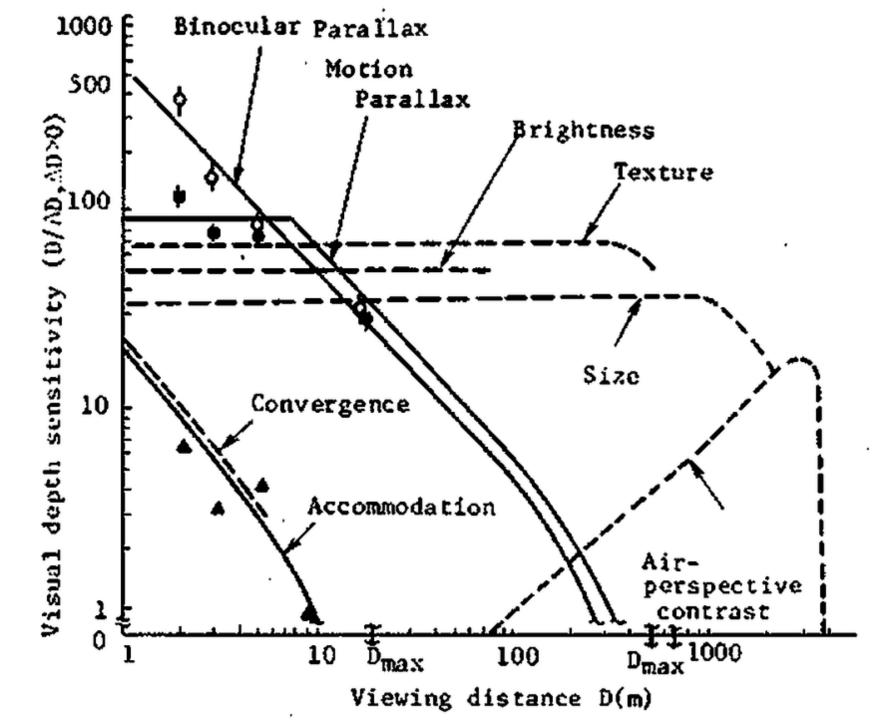


## This is how 3D movies work



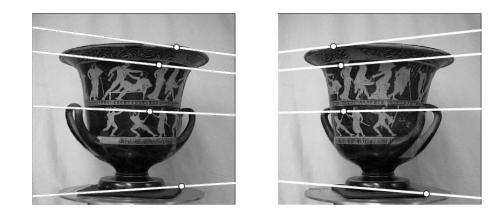


Is disparity the only depth cue the human visual system uses?

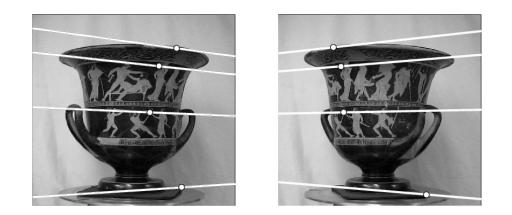


Nagata '89

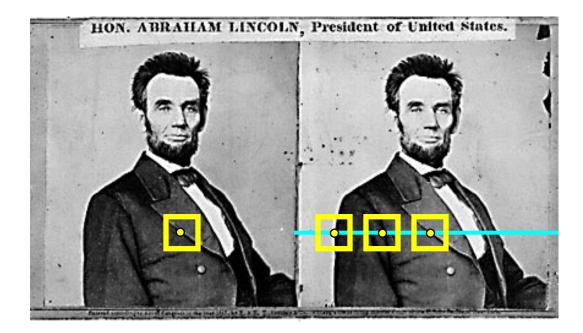
# So can I compute depth from any two images of the same object?



## So can I compute depth from any two images of the same object?



- 1. Need sufficient baseline
- 2. Images need to be 'rectified' first (make epipolar lines horizontal)

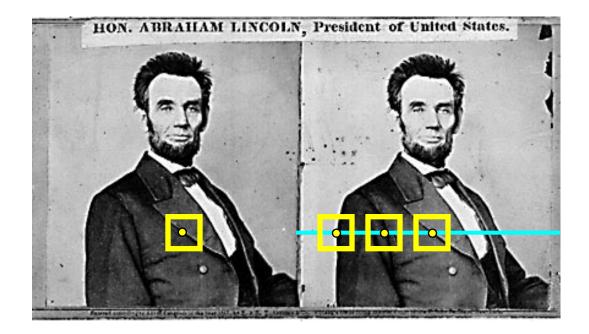


#### 1. Rectify images

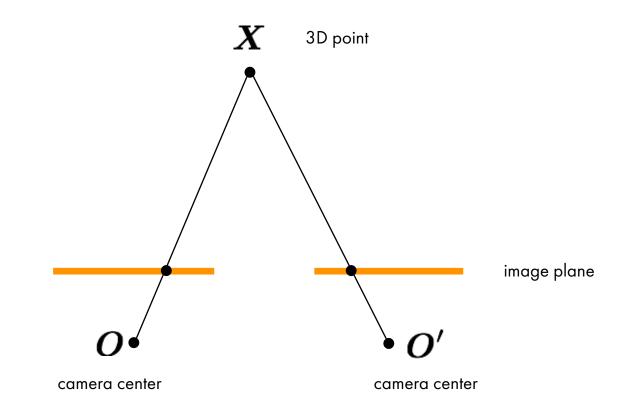
(make epipolar lines horizontal)

- 2. For each pixel
  - a. Find epipolar line
  - b.Scan line for best match
  - c.Compute depth from disparity

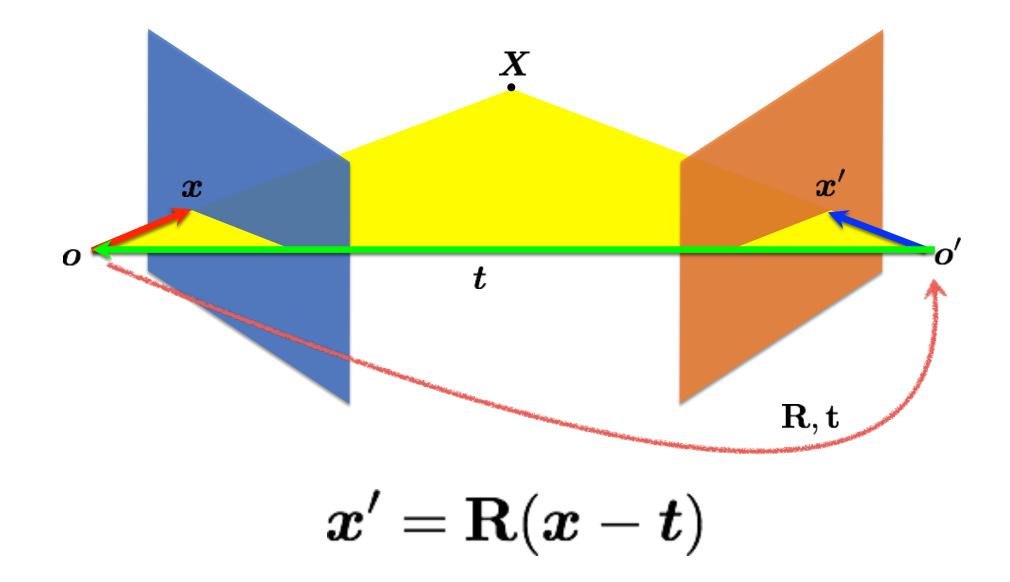
$$Z = \frac{bf}{d}$$



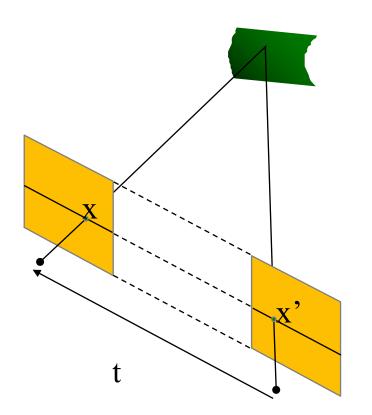
### How can you make the epipolar lines horizontal?



#### What's special about these two cameras?

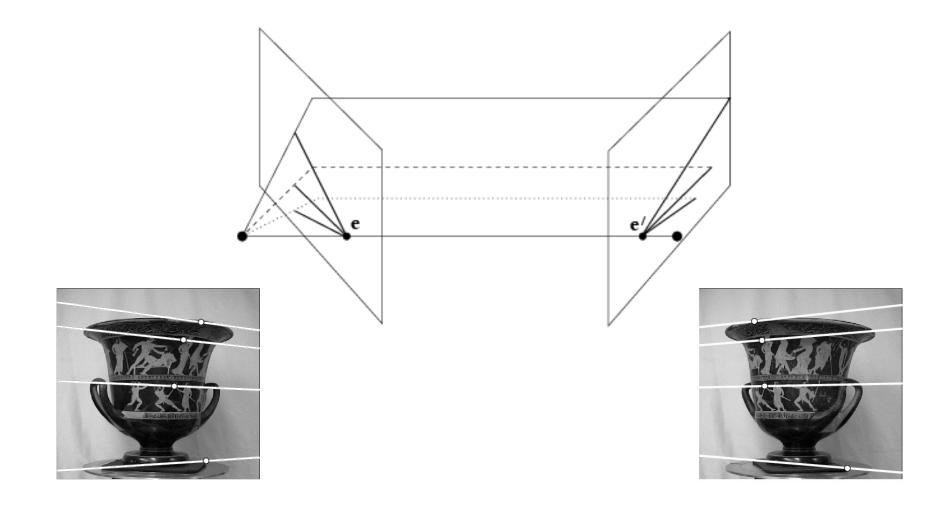


## When are epipolar lines horizontal?



When this relationship holds:

$$R = I \qquad t = (T, 0, 0)$$



## It's hard to make the image planes exactly parallel



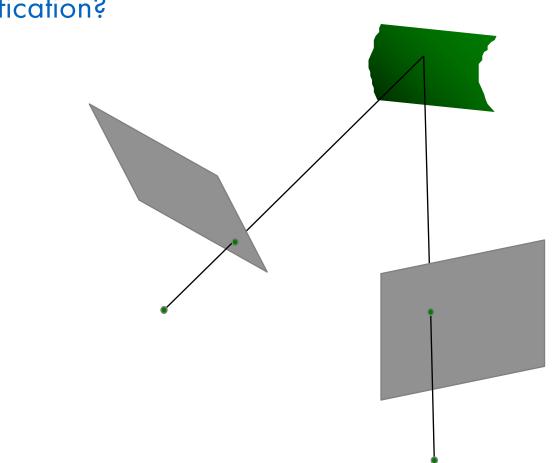
## How can you make the epipolar lines horizontal?





#### Use stereo rectification?

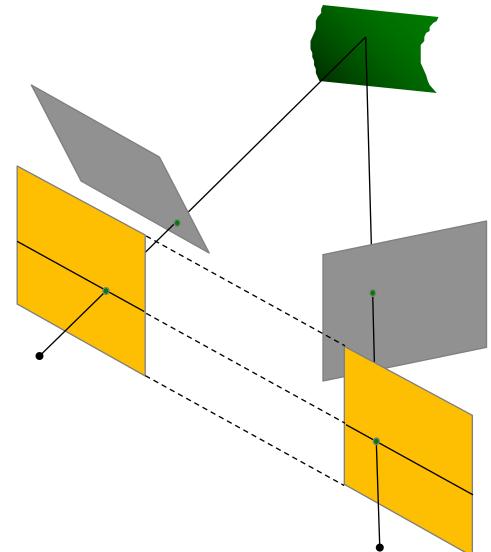




#### What is stereo rectification?

#### What is stereo rectification?

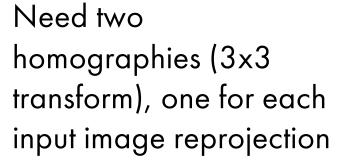
Reproject image planes onto a common plane parallel to the line between camera centers

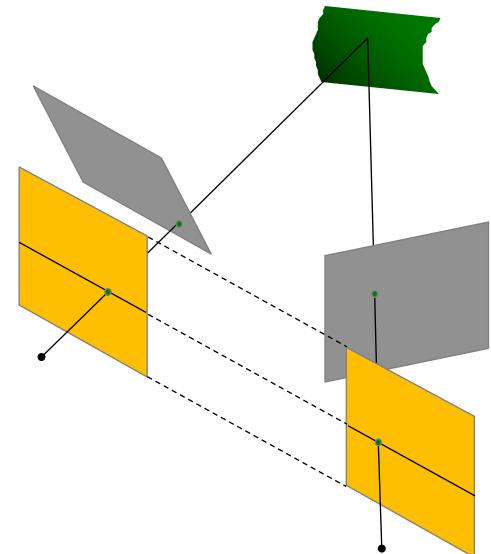


How can you do this?

#### What is stereo rectification?

Reproject image planes onto a common plane parallel to the line between camera centers

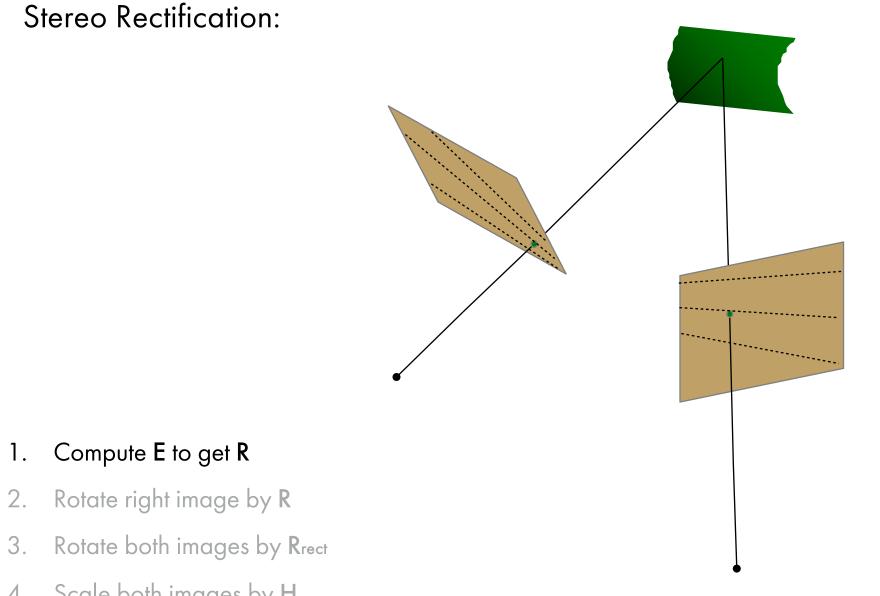




C. Loop and Z. Zhang. Computing Rectifying Homographies for Stereo Vision. Computer Vision and Pattern Recognition, 1999.

- Rotate the right camera by R

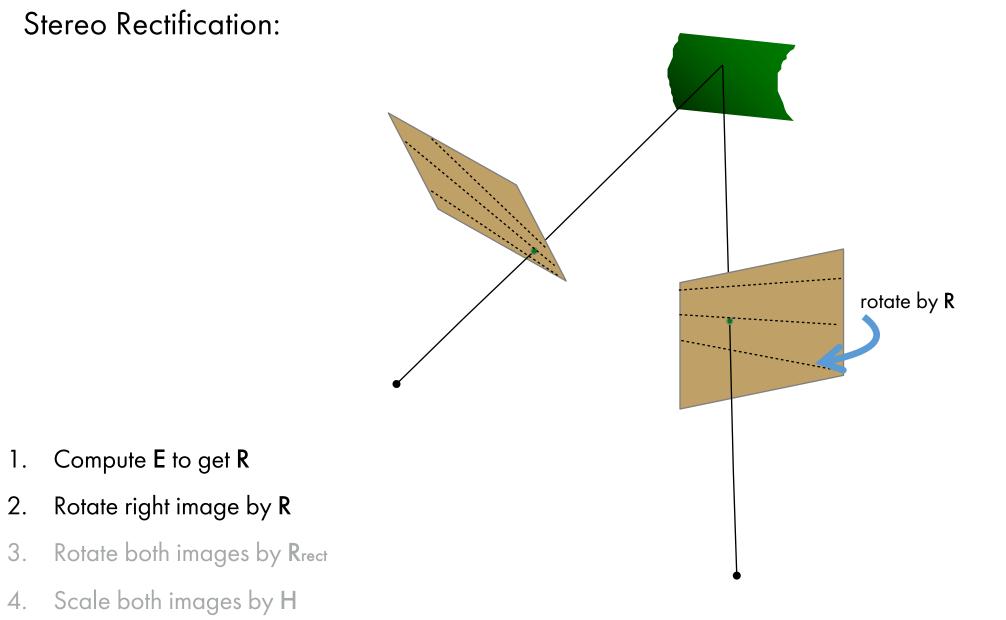
   (aligns camera coordinate system orientation only)
- 2. Rotate (rectify) the left camera so that the epipole is at infinity
- 3. Rotate (rectify) the right camera so that the epipole is at infinity
- 4. Adjust the scale



4. Scale both images by H

2. Rotate right image by **R** 

1. Compute **E** to get **R** 



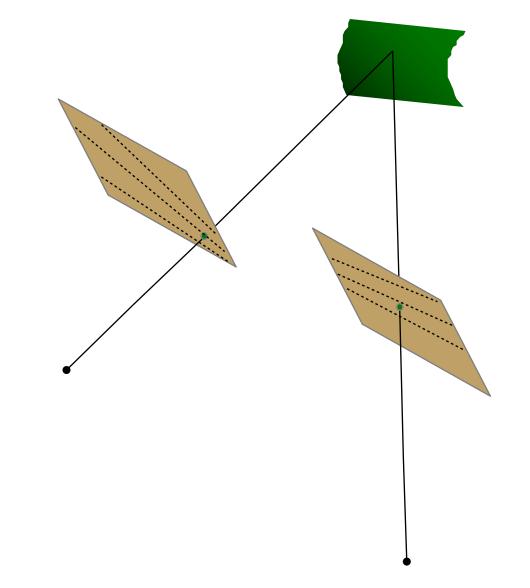
4. Scale both images by H

Compute **E** to get **R** 

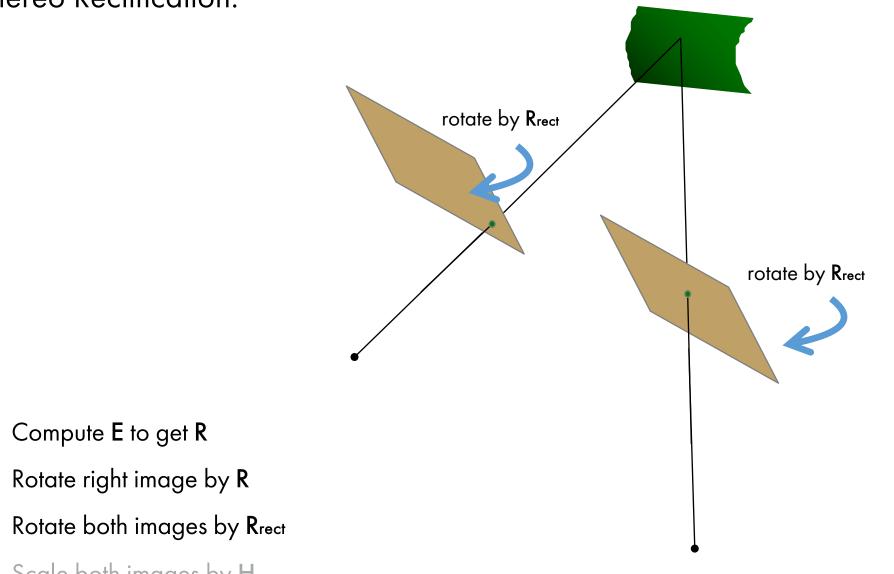
Rotate right image by **R** 

1.

2.



- 1. Compute **E** to get **R**
- 2. Rotate right image by **R**
- 3. Rotate both images by **R**rect
- 4. Scale both images by **H**



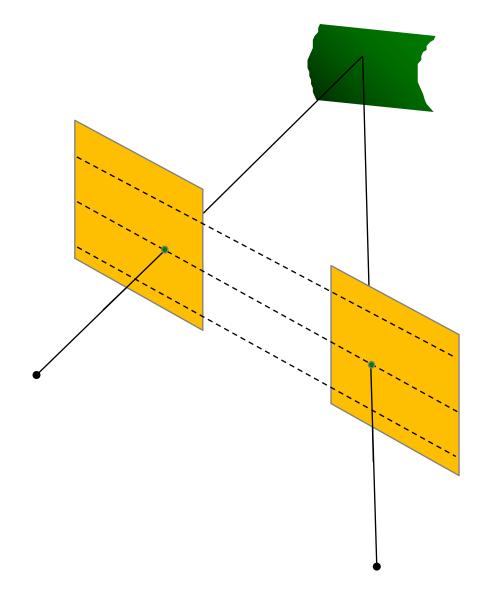
Rotate both images by **R**rect 3.

Compute E to get R

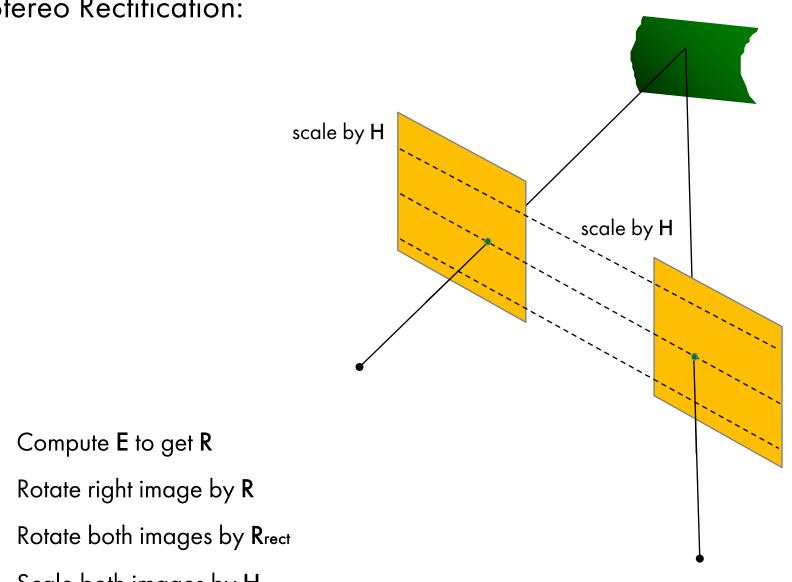
1.

2.

Scale both images by **H** 4.



- 1. Compute **E** to get **R**
- 2. Rotate right image by **R**
- 3. Rotate both images by **R**<sub>rect</sub>
- 4. Scale both images by **H**



Scale both images by **H** 4.

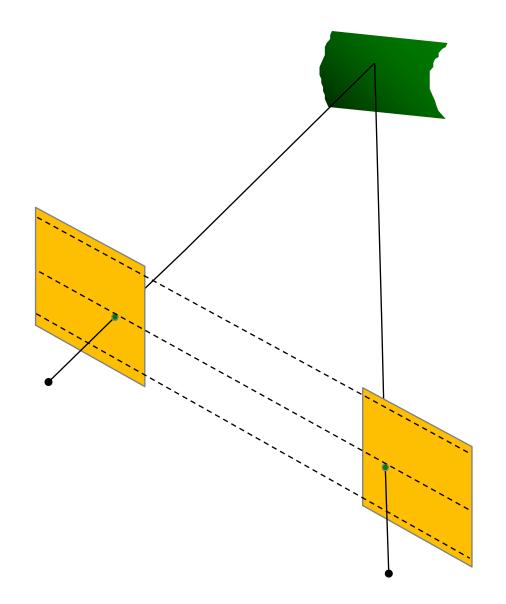
Compute E to get R

Rotate right image by **R** 

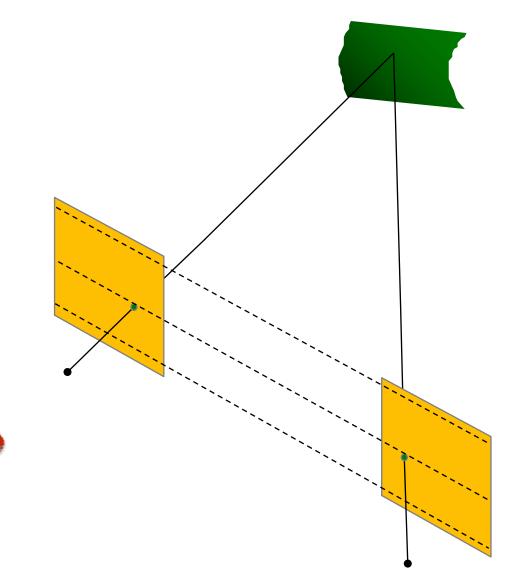
1.

2.

3.



- 1. Compute **E** to get **R**
- 2. Rotate right image by **R**
- 3. Rotate both images by **R**<sub>rect</sub>
- 4. Scale both images by **H**



- 1. Compute **E** to get **R**
- 2. Rotate right image by **R**
- 3. Rotate both images by **R**<sub>rect</sub>
- 4. Scale both images by **H**

# Step 1: Compute E to get R

SVD: 
$$\mathbf{E} = \mathbf{U} \mathbf{\Sigma} \mathbf{V}^{\top}$$
 Let  $\mathbf{W} = \begin{bmatrix} 0 & -1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 1 \end{bmatrix}$ 

We get FOUR solutions:

$$\mathbf{E} = [\mathbf{R}|\mathbf{T}]$$

$$\mathbf{R}_1 = \mathbf{U}\mathbf{W}\mathbf{V}^{\top}$$
  $\mathbf{R}_2 = \mathbf{U}\mathbf{W}^{\top}\mathbf{V}^{\top}$   $\mathbf{T}_1 = U_3$   $\mathbf{T}_2 = -U_3$ 

two possible rotations

two possible translations

We get FOUR solutions:

$$\mathbf{R}_1 = \mathbf{U}\mathbf{W}\mathbf{V}^{ op}$$
  $\mathbf{R}_1 = \mathbf{U}\mathbf{W}\mathbf{V}^{ op}$   
 $\mathbf{T}_1 = U_3$   $\mathbf{T}_2 = -U_3$ 

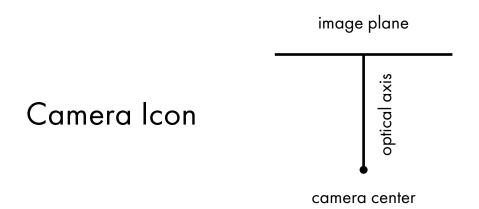
 $\mathbf{R}_2 = \mathbf{U}\mathbf{W}^\top \mathbf{V}^\top \qquad \qquad \mathbf{R}_2 = \mathbf{U}\mathbf{W}^\top \mathbf{V}^\top \\ \mathbf{T}_2 = -U_3 \qquad \qquad \qquad \mathbf{T}_1 = U_3$ 

#### Which one do we choose?

Compute determinant of R, valid solution must be equal to 1 (note: det(R) = -1 means rotation and reflection)

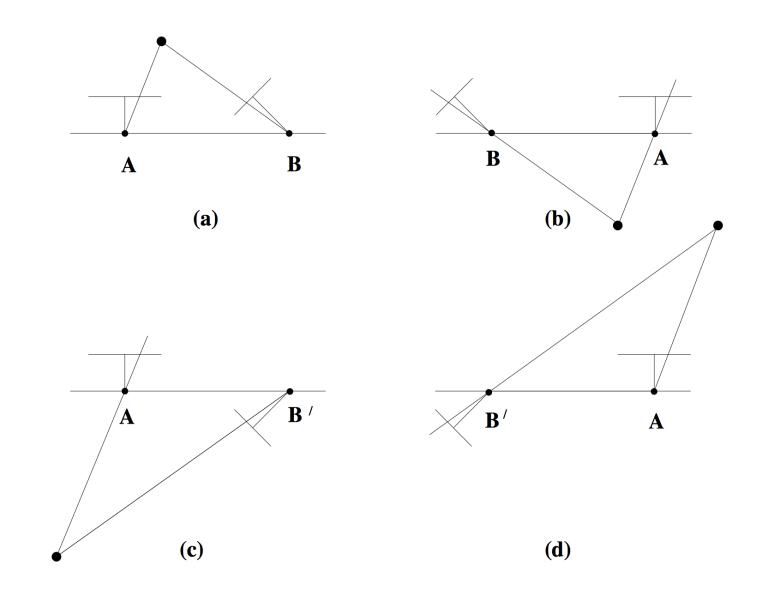
Compute 3D point using triangulation, valid solution has positive Z value (Note: negative Z means point is behind the camera )

#### Let's visualize the four configurations...

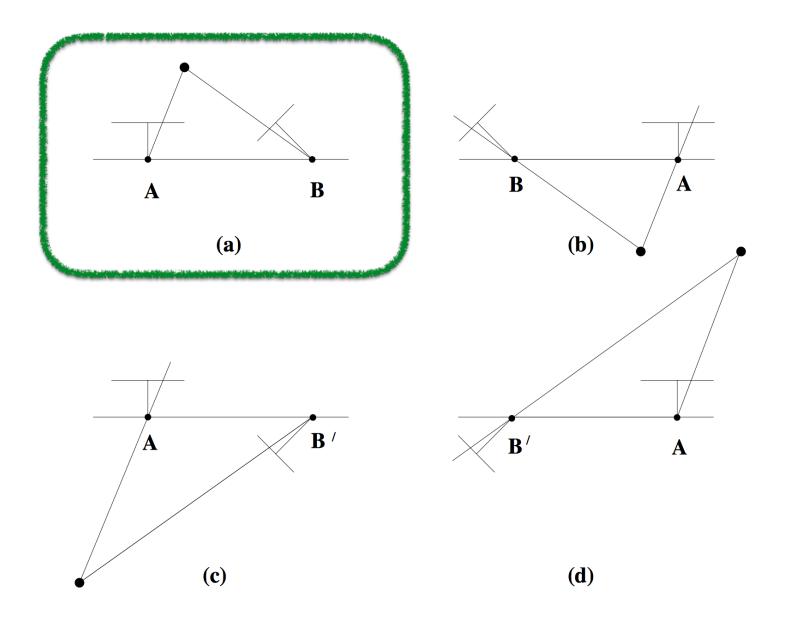


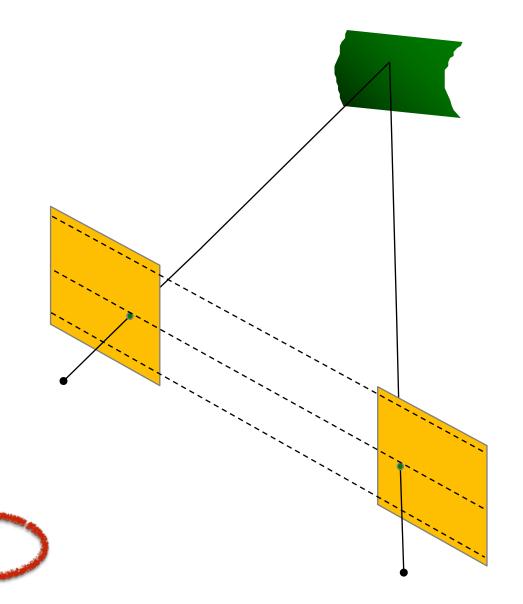
#### Find the configuration where the point is in front of both cameras

Find the configuration where the point is in front of both cameras



Find the configuration where the points is in front of both cameras

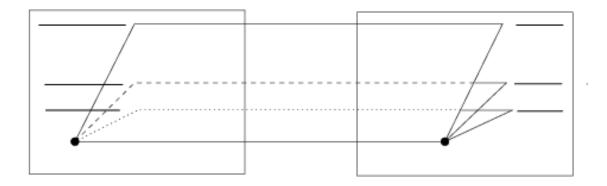




- 1. Compute E to get R
- 2. Rotate right image by R
- 3. Rotate both images by Rrect
- 4. Scale both images by H

When do epipolar lines become horizontal?

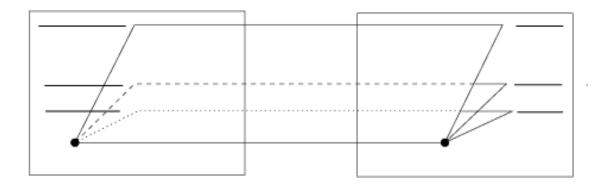
#### Parallel cameras

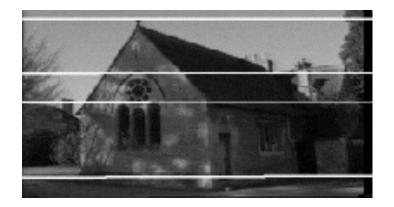


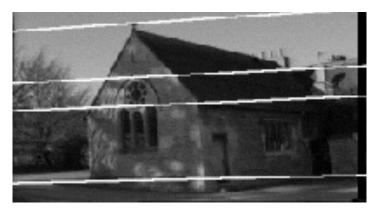


#### Where is the epipole?

#### Parallel cameras







epipole at infinity

# Setting the epipole to infinity (Building Rrect from E)

 $\boldsymbol{r}_3 = \boldsymbol{r}_1 imes \boldsymbol{r}_2$ 

Let 
$$R_{\text{rect}} = \begin{bmatrix} \boldsymbol{r}_1^\top \\ \boldsymbol{r}_2^\top \\ \boldsymbol{r}_3^\top \end{bmatrix}$$
 Given: (using SVD on E)  
(translation from E)

$$m{r}_1 = m{e}_1 = rac{T}{||T||}$$
 epipole coincides with translation vector $m{r}_2 = rac{1}{\sqrt{T_x^2 + T_y^2}} \begin{bmatrix} -T_y & T_x & 0 \end{bmatrix}$  cross product of e and the direction vector of the optical axis

orthogonal vector

If 
$$\mathbf{r}_1 = \mathbf{e}_1 = \frac{T}{||T||}$$
 and  $\mathbf{r}_2$   $\mathbf{r}_3$  orthogona  
then  $R_{\text{rect}}\mathbf{e}_1 = \begin{bmatrix} \mathbf{r}_1^{\top}\mathbf{e}_1 \\ \mathbf{r}_2^{\top}\mathbf{e}_1 \\ \mathbf{r}_3^{\top}\mathbf{e}_1 \end{bmatrix} = \begin{bmatrix} ? \\ ? \\ ? \end{bmatrix}$ 

If 
$$r_1 = e_1 = \frac{T}{||T||}$$
 and  $r_2$   $r_3$  orthogonal  
then  $R_{\text{rect}}e_1 = \begin{bmatrix} r_1^\top e_1 \\ r_2^\top e_1 \\ r_3^\top e_1 \end{bmatrix} = \begin{bmatrix} 1 \\ 0 \\ 0 \end{bmatrix}$ 

Where is this point located on the image plane?

If 
$$r_1 = e_1 = \frac{T}{||T||}$$
 and  $r_2$   $r_3$  orthogonal  
then  $R_{\text{rect}}e_1 = \begin{bmatrix} r_1^\top e_1 \\ r_2^\top e_1 \\ r_3^\top e_1 \end{bmatrix} = \begin{bmatrix} 1 \\ 0 \\ 0 \end{bmatrix}$ 

Where is this point located on the image plane?

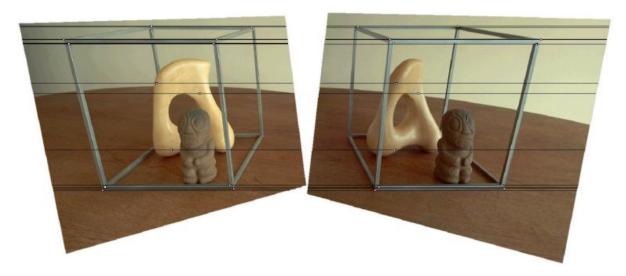
At x-infinity

#### Stereo Rectification Algorithm

- 1. Estimate E using the 8 point algorithm (SVD)
- 2. Estimate the epipole e (SVD of E)
- 3. Build **R**rect from **e**
- 4. Decompose E into R and T
- 5. Set  $\mathbf{R}_1 = \mathbf{R}_{\text{rect}}$  and  $\mathbf{R}_2 = \mathbf{R}\mathbf{R}_{\text{rect}}$
- 6. Rotate each left camera point (warp image)
   [x' y' z'] = R1 [x y z]
- 7. Rectified points as  $\mathbf{p} = f/z'[x' y' z']$
- 8. Repeat 6 and 7 for right camera points using  $\mathbf{R}_2$



# What can we do after rectification?

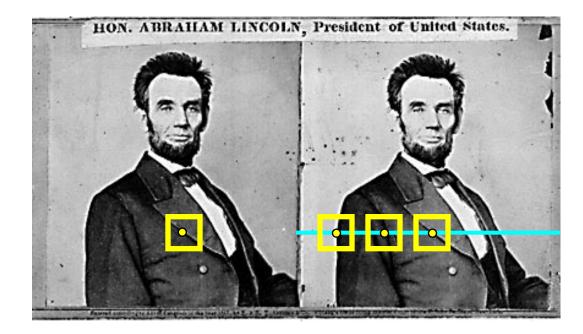


# Stereo matching



## Depth Estimation via Stereo Matching





#### 1. Rectify images

(make epipolar lines horizontal)

#### 2.For each pixel a.Find epipolar line

b.Scan line for best match .

c.Compute depth from disparity

$$Z = \frac{bf}{d}$$

How would you do this?

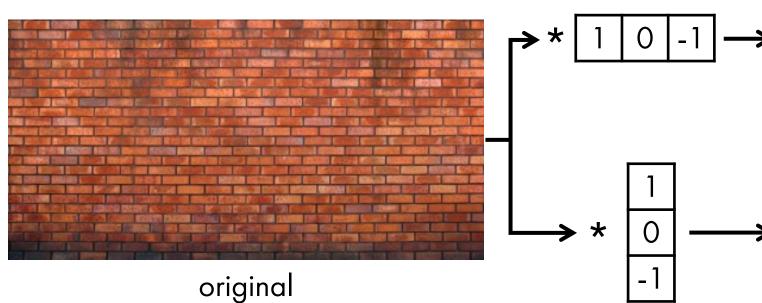
### Reminder from filtering

How do we detect an edge?

### Reminder from filtering

How do we detect an edge?

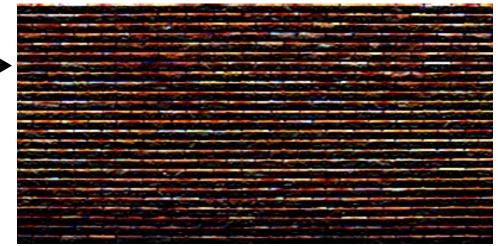
• We filter with something that looks like an edge.



We can think of linear filtering as a way to evaluate how similar an image is locally to some template.

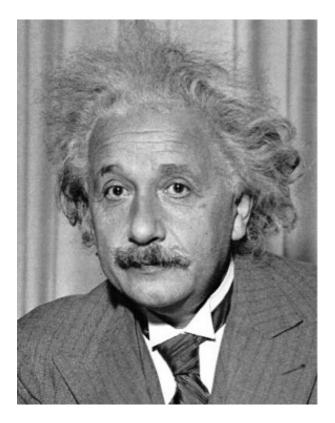


#### horizontal edge filter

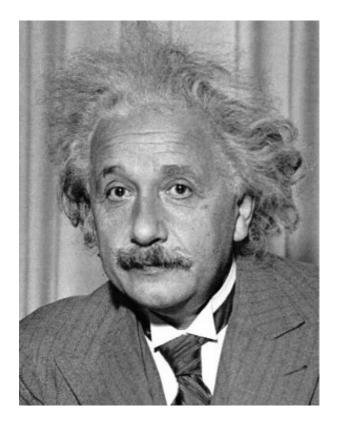


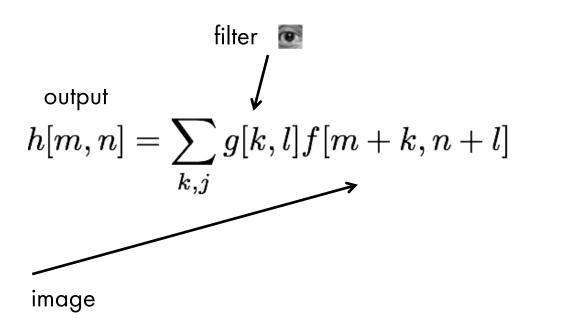
vertical edge filter

How do we detect the template I mage?



How do we detect the template I mage?

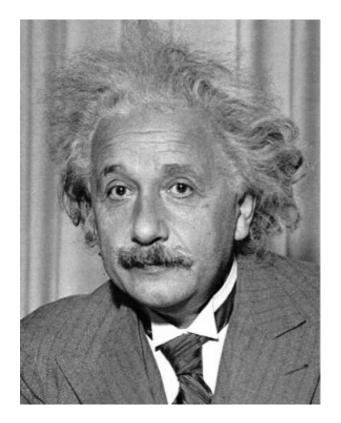


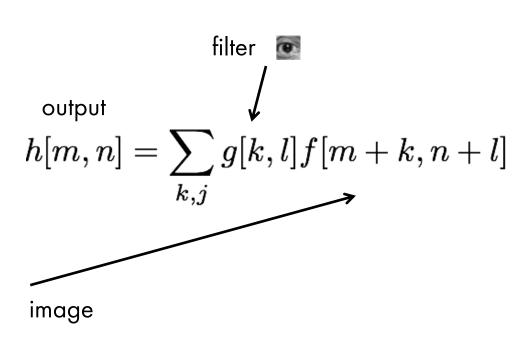


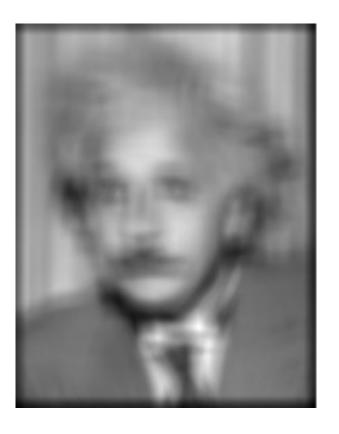
What will the output look like?

Solution 1: Filter the image using the template as filter kernel.

How do we detect the template I mage?



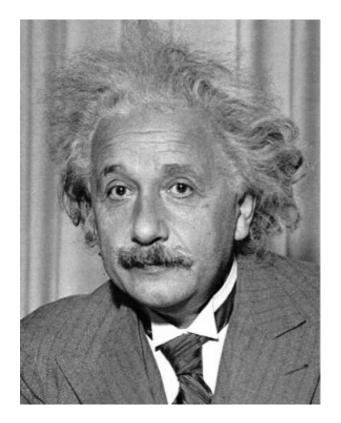


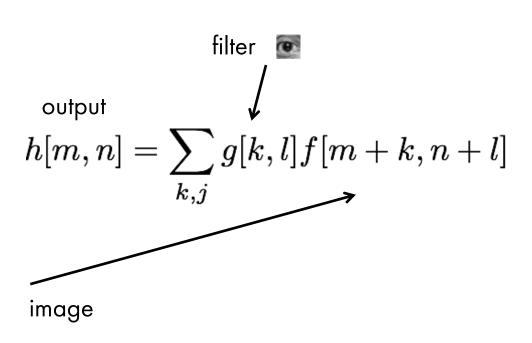


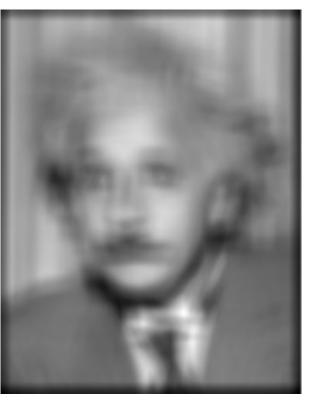
Solution 1: Filter the image using the template as filter kernel.

What went wrong?

How do we detect the template I mage?



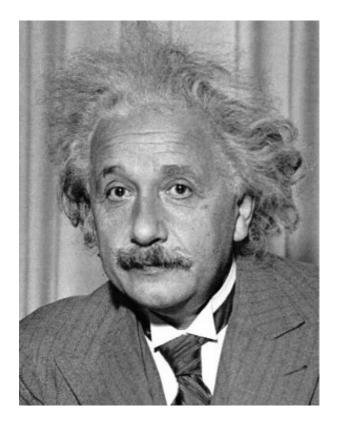


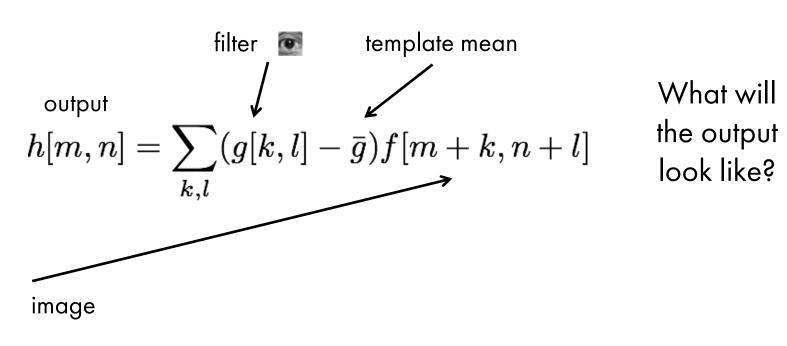


Increases for higher local intensities.

Solution 1: Filter the image using the template as filter kernel.

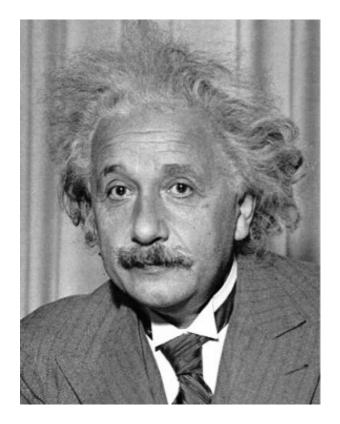
How do we detect the template I mage?

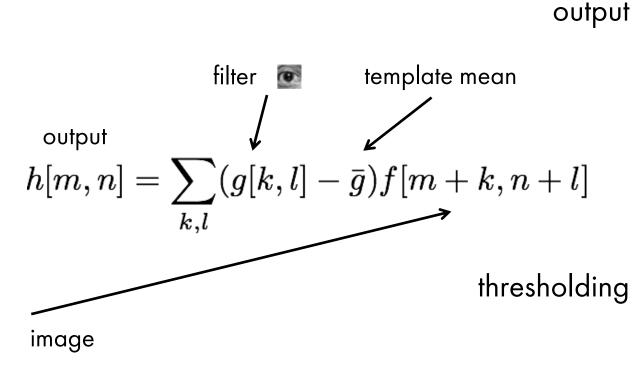


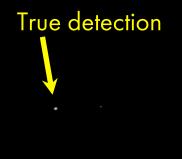


Solution 2: Filter the image using a zero-mean template.

How do we detect the template I mage?





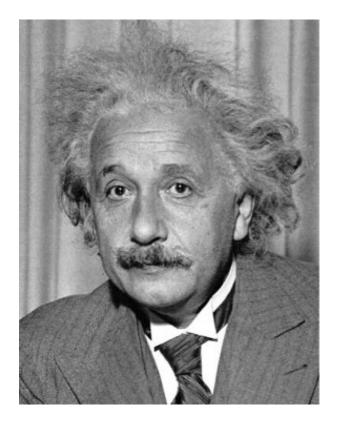


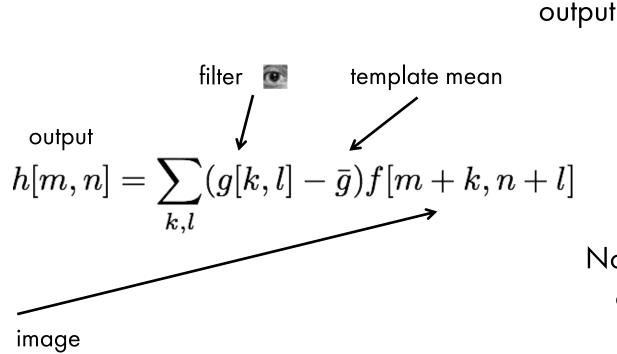


What went wrong?

Solution 2: Filter the image using a zero-mean template.

How do we detect the template I mage?

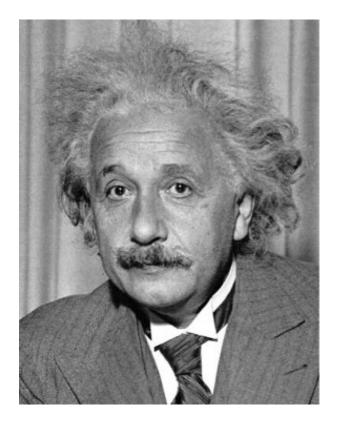


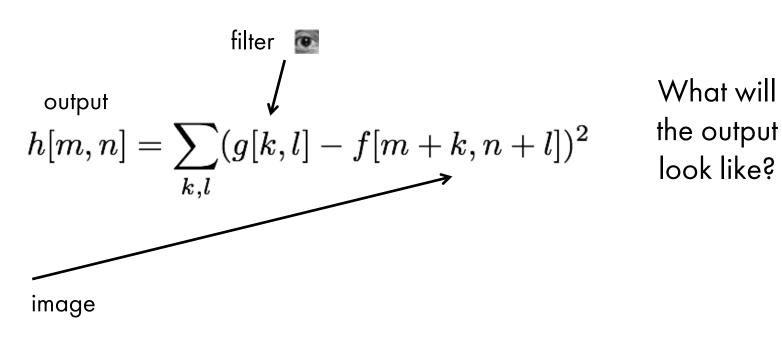


Not robust to highcontrast areas

Solution 2: Filter the image using a zero-mean template.

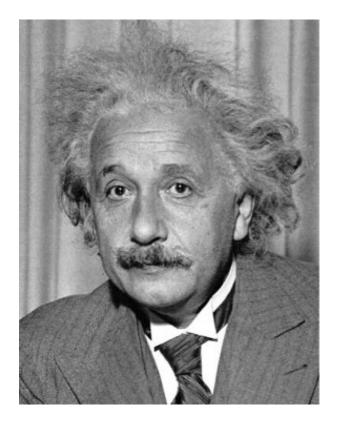
How do we detect the template I mage?

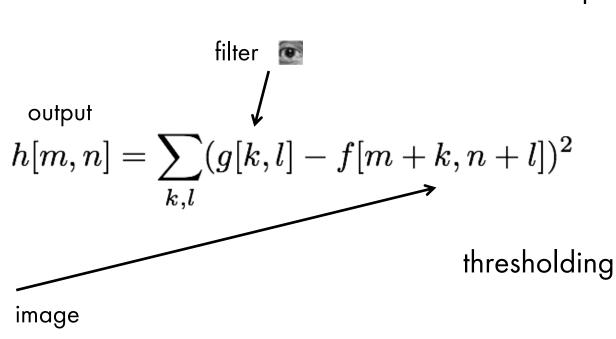




Solution 3: Use sum of squared differences (SSD).

How do we detect the template I mage?





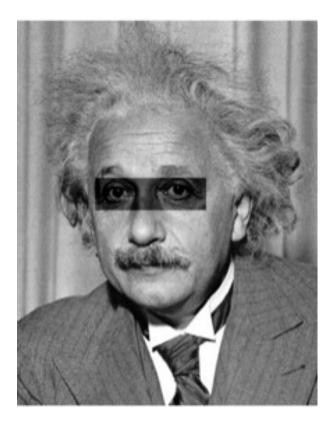
1-output

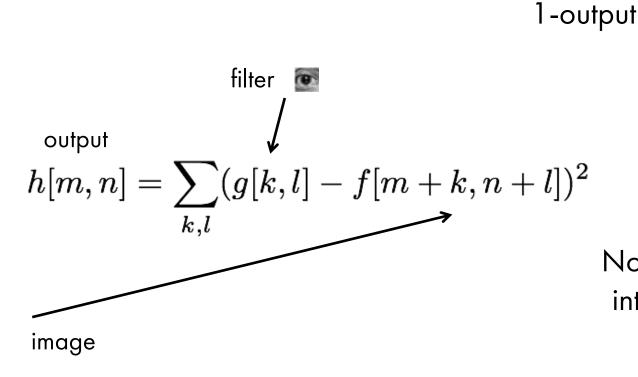
True detection

Solution 3: Use sum of squared differences (SSD).

What could go wrong?

How do we detect the template I mage?

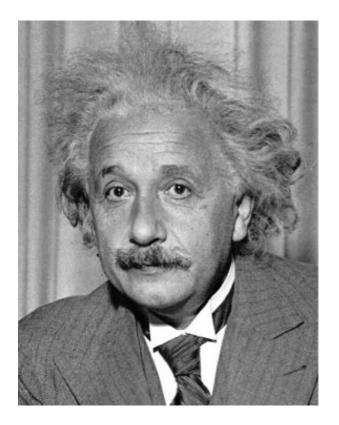




Not robust to local intensity changes

Solution 3: Use sum of squared differences (SSD).

How do we detect the template I mage?

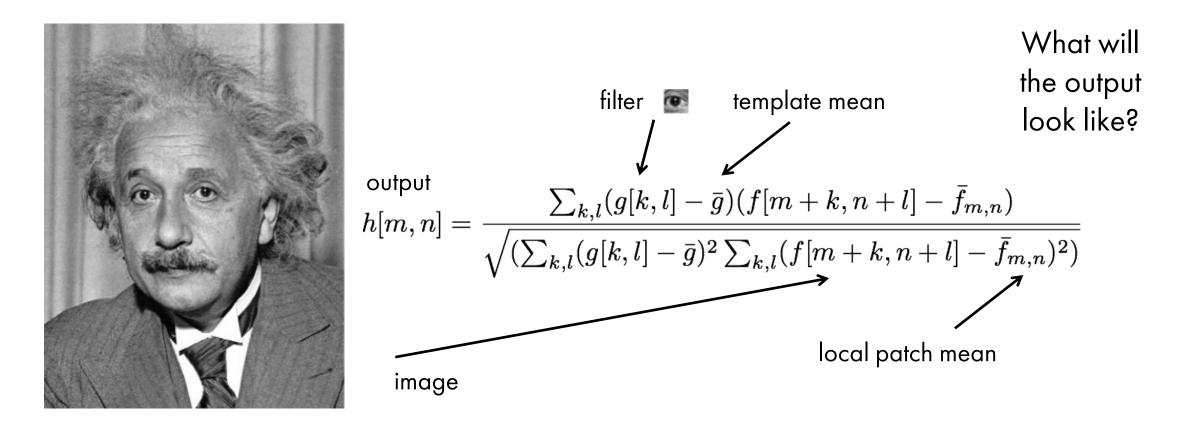


Observations so far:

- subtracting mean deals with brightness bias
- dividing by standard deviation removes contrast bias

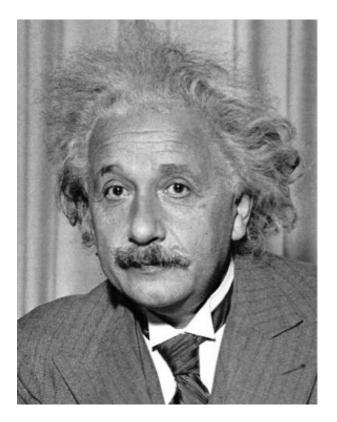
Can we combine the two effects?

How do we detect the template I mage?



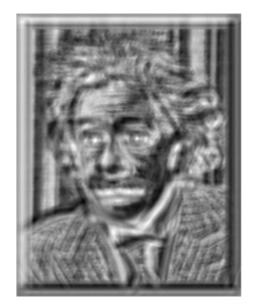
Solution 4: Normalized cross-correlation (NCC).

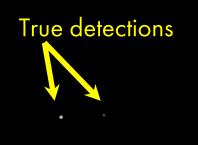
How do we detect the template I mage?



Solution 4: Normalized cross-correlation (NCC).

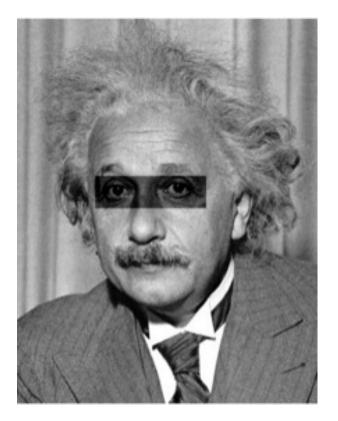
1-output





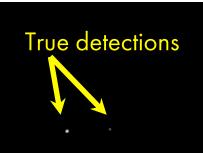
thresholding

How do we detect the template I mage?



Solution 4: Normalized cross-correlation (NCC).

1-output



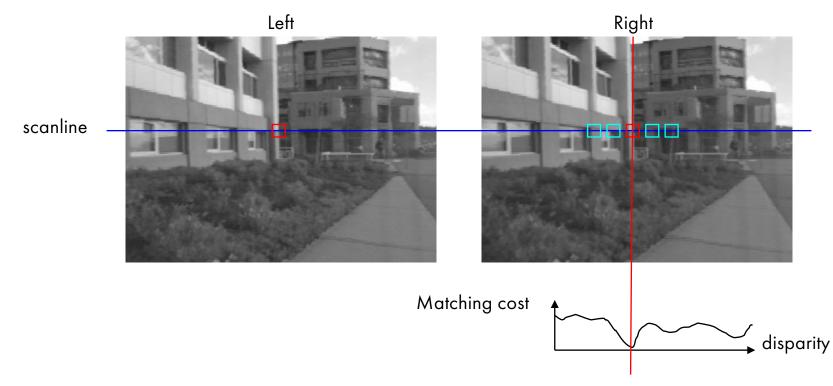
thresholding

#### What is the best method?

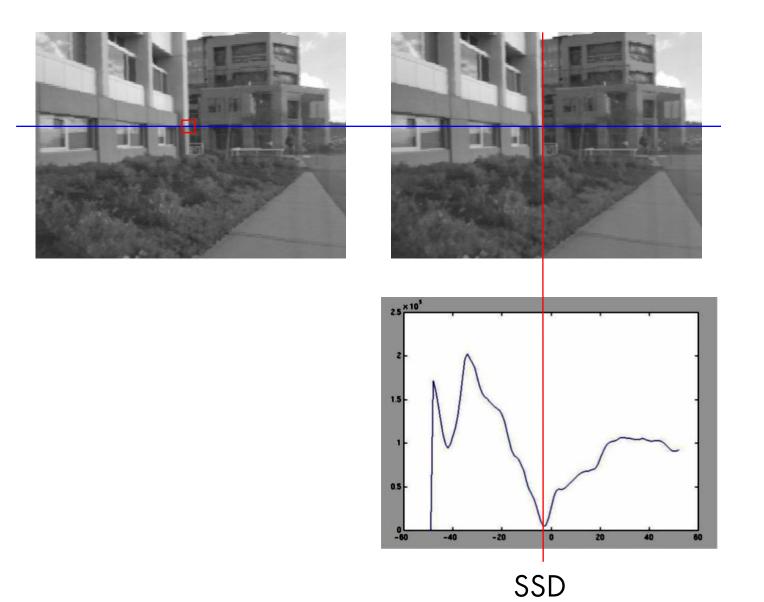
It depends on whether you care about speed or invariance.

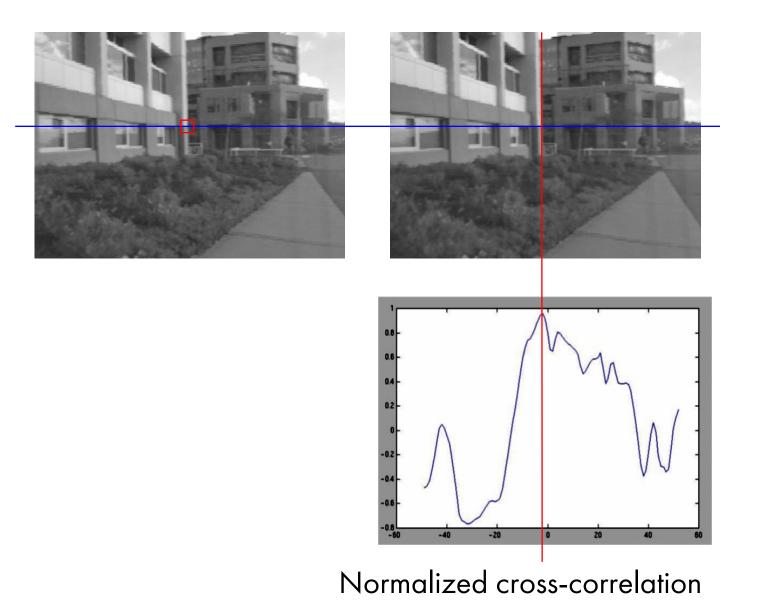
- Zero-mean: Fastest, very sensitive to local intensity.
- Sum of squared differences: Medium speed, sensitive to intensity offsets.
- Normalized cross-correlation: Slowest, invariant to contrast and brightness.

# Stereo Block Matching

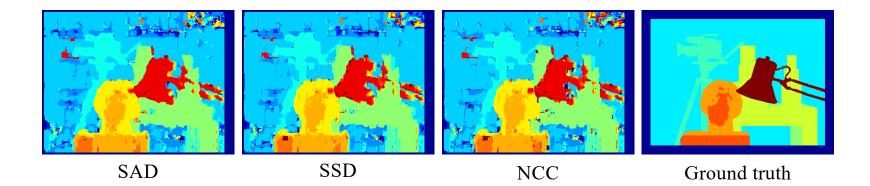


- Slide a window along the epipolar line and compare contents of that window with the reference window in the left image
- Matching cost: SSD or normalized correlation





Similarity Measure	Formula
Sum of Absolute Differences (SAD)	$\sum_{(i,j)\in W}  I_1(i,j) - I_2(x+i,y+j) $
Sum of Squared Differences (SSD)	$\sum_{(i,j)\in W} \left( I_1(i,j) - I_2(x+i,y+j) \right)^2$
Zero-mean SAD	$\sum_{(i,j)\in W}  I_1(i,j) - \bar{I}_1(i,j) - I_2(x+i,y+j) + \bar{I}_2(x+i,y+j) $ $\bar{I}_2(x+i,y+j) = \bar{I}_2(x+i,y+j) - \bar{I}_2(x+i,y+j) = \bar{I}_2(x+$
Locally scaled SAD	$\sum_{(i,j)\in W}  I_1(i,j) - \frac{\bar{I}_1(i,j)}{\bar{I}_2(x+i,y+j)} I_2(x+i,y+j) $ $\sum_{(i,j)\in W} I_1(i,j) . I_2(x+i,y+j)$
Normalized Cross Correlation (NCC)	$\frac{\sum_{(i,j)\in W} I_1(i,j) \cdot I_2(x+i,y+j)}{\sqrt[2]{\sum_{(i,j)\in W} I_1^2(i,j) \cdot \sum_{(i,j)\in W} I_2^2(x+i,y+j)}}$



#### Effect of window size



W = 3

W = 20

#### Effect of window size







W = 3

W = 20

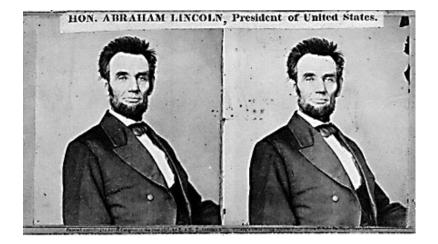
#### Smaller window

- + More detail
- More noise

#### Larger window

- + Smoother disparity maps
- Less detail
- Fails near boundaries

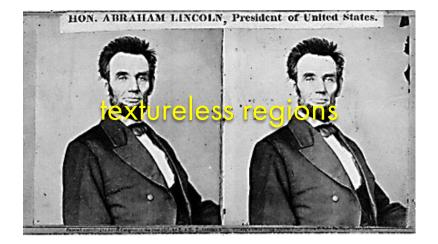
#### When will stereo block matching fail?







#### When will stereo block matching fail?





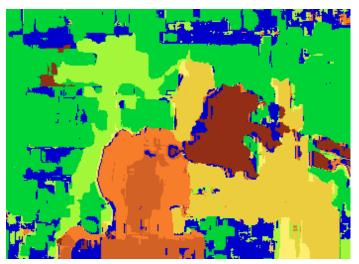


#### Improving stereo matching



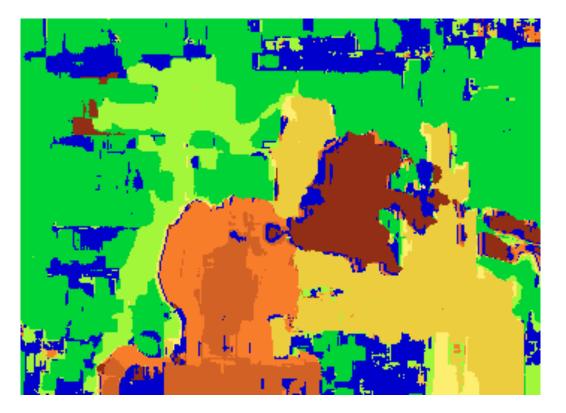
#### Block matching

#### Ground truth

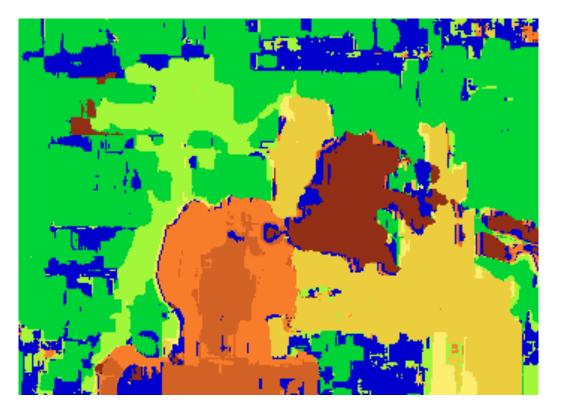




What are some problems with the result?



How can we improve depth estimation?



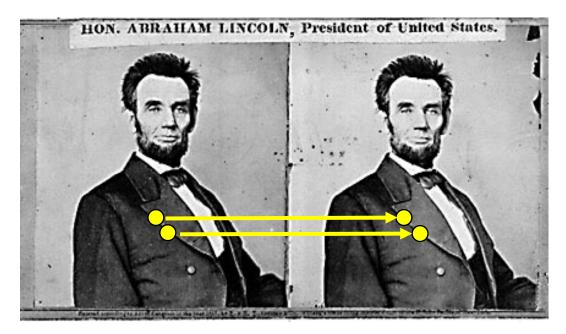
#### How can we improve depth estimation?

Too many discontinuities. We expect disparity values to change slowly.

> Let's make an assumption: depth should change smoothly

Stereo matching as ...

# Energy Minimization



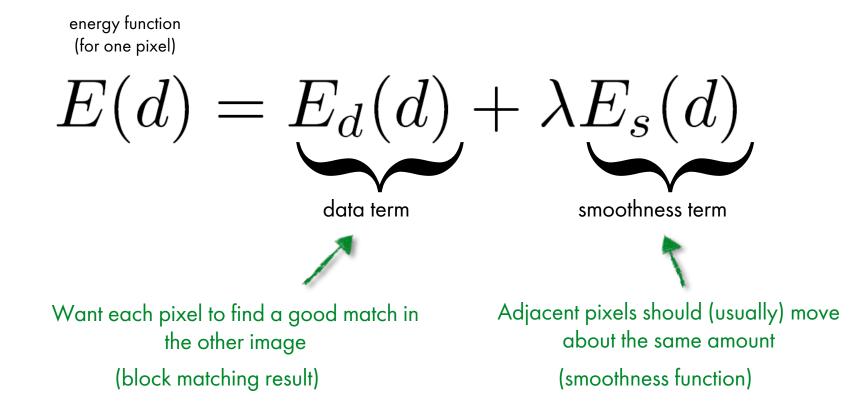
What defines a good stereo correspondence?

#### 1. Match quality

- Want each pixel to find a good match in the other image

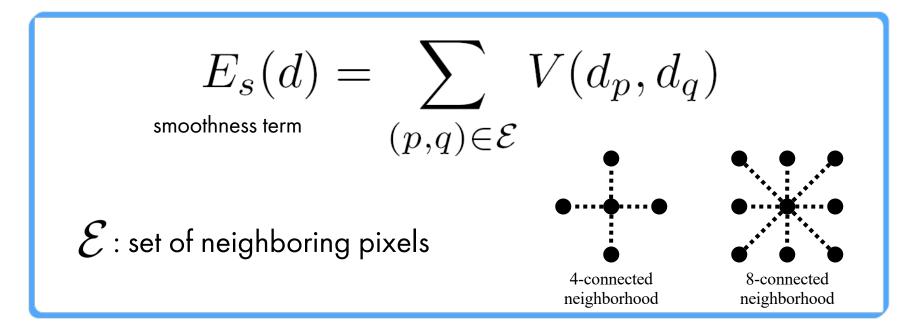
#### 2. Smoothness

If two pixels are adjacent, they should (usually) move about the same amount



$$\begin{split} E(d) = E_d(d) + \lambda E_s(d) \\ E_d(d) = \sum_{(x,y) \in I} C(x,y,d(x,y)) \\ & \text{data term} \end{split} \text{SSD distance between windows centered} \\ & \text{distance between windows c$$

 $E(d) = E_d(d) + \lambda E_s(d)$  $E_d(d) = \sum C(x, y, d(x, y))$  $(x,y) \in I$ SSD distance between windows centered at I(x, y) and J(x+d(x,y), y)



$$E_s(d) = \sum_{(p,q) \in \mathcal{E}} V(d_p,d_q)$$
 smoothness term

$$V(d_p, d_q) = |d_p - d_q|$$

One possible solution...

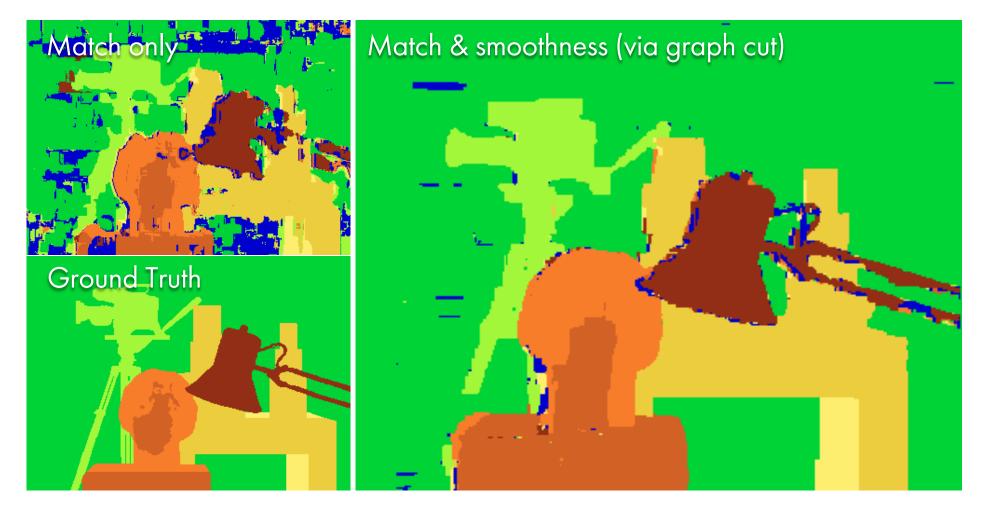
#### Dynamic Programming

$$E(d) = E_d(d) + \lambda E_s(d)$$

Can minimize this independently per scanline using dynamic programming (DP)......

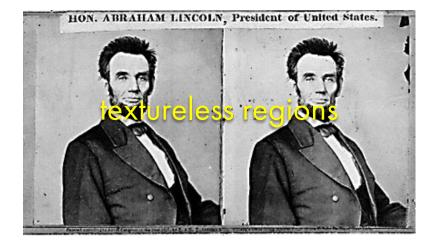
D(x,y,d) : minimum cost of solution such that d(x,y) = d

$$D(x, y, d) = C(x, y, d) + \min_{d'} \left\{ D(x - 1, y, d') + \lambda \left| d - d' \right| \right\}$$



Y. Boykov, O. Veksler, and R. Zabih, Fast Approximate Energy Minimization via Graph Cuts, PAMI 2001

#### All of these cases remain difficult, what can we do?



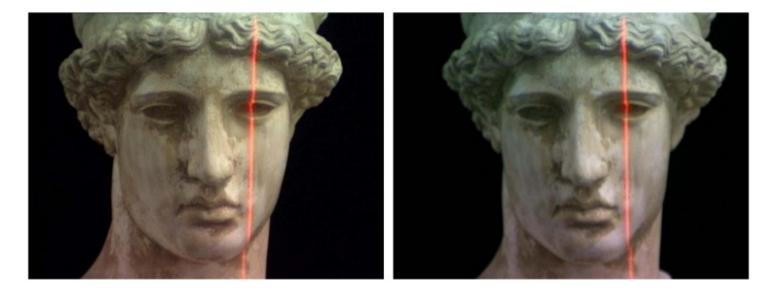


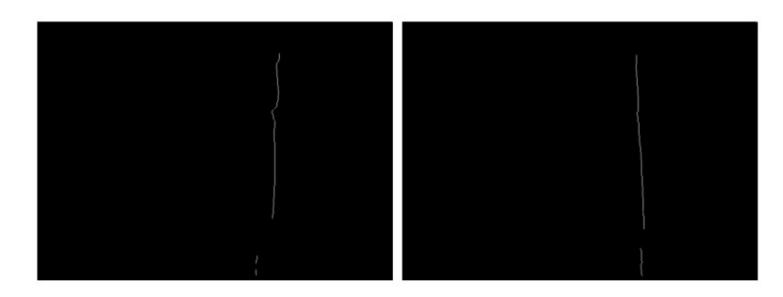


## Structured light

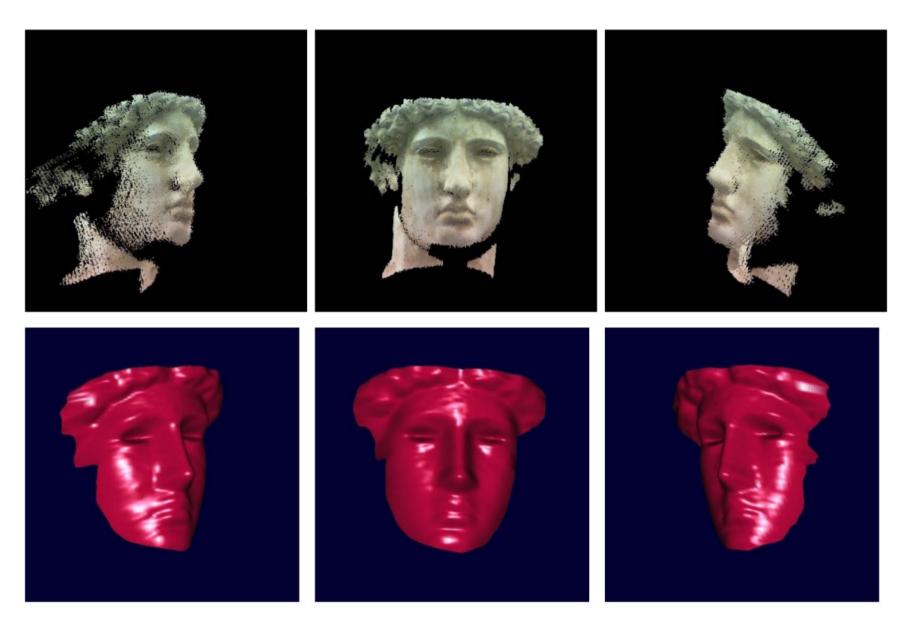
### Use controlled ("structured") light to make correspondences easier

Disparity between laser points on the same scanline in the images determines the 3-D coordinates of the laser point on object

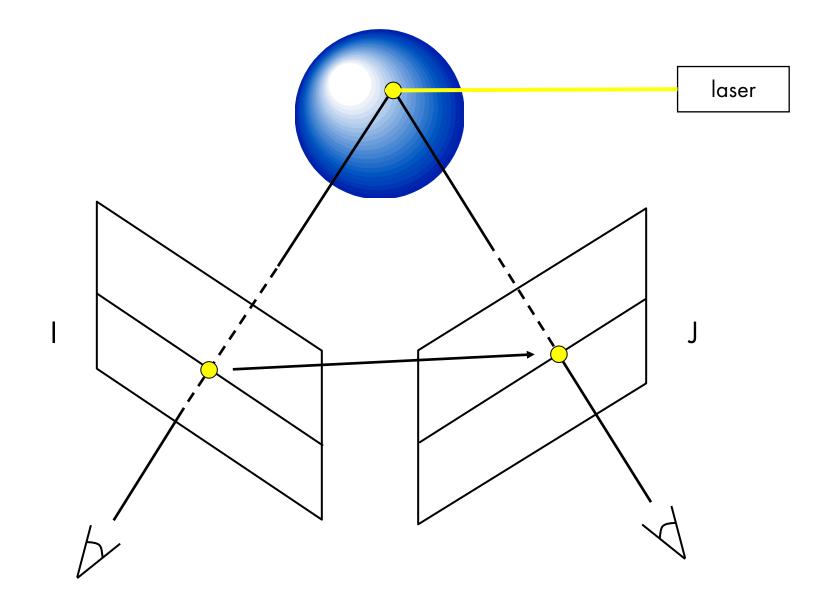




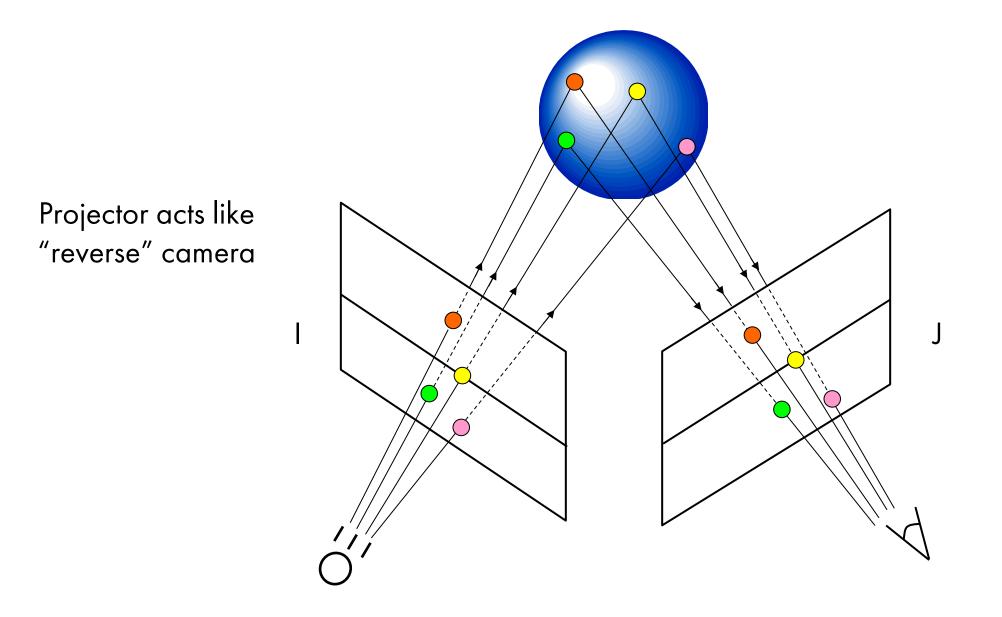
### Use controlled ("structured") light to make correspondences easier



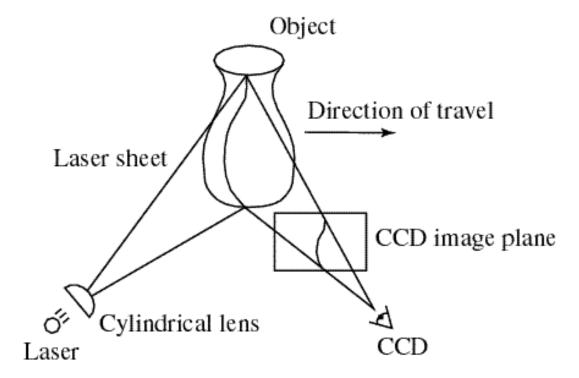
## Structured light and two cameras

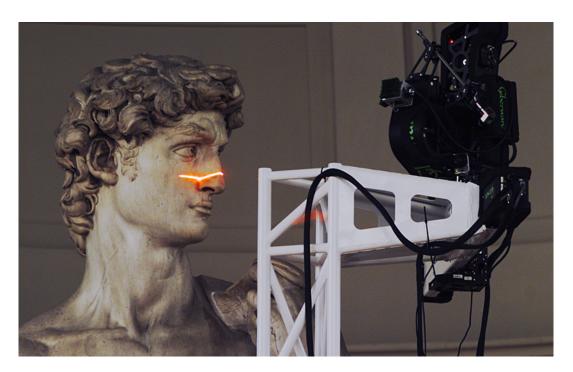


### Structured light and one camera



### Example: Laser scanner





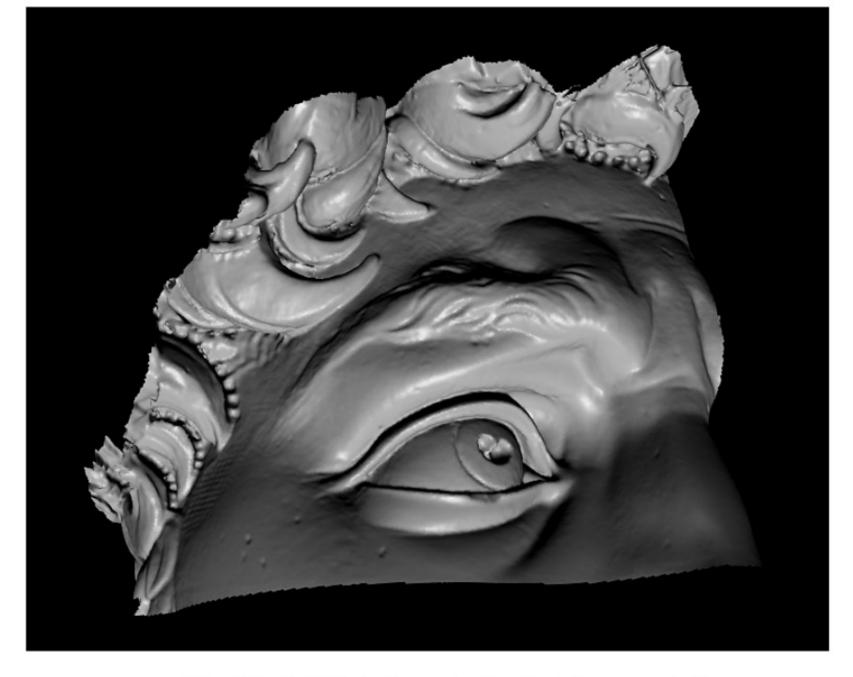
Digital Michelangelo Project http://graphics.stanford.edu/projects/mich/



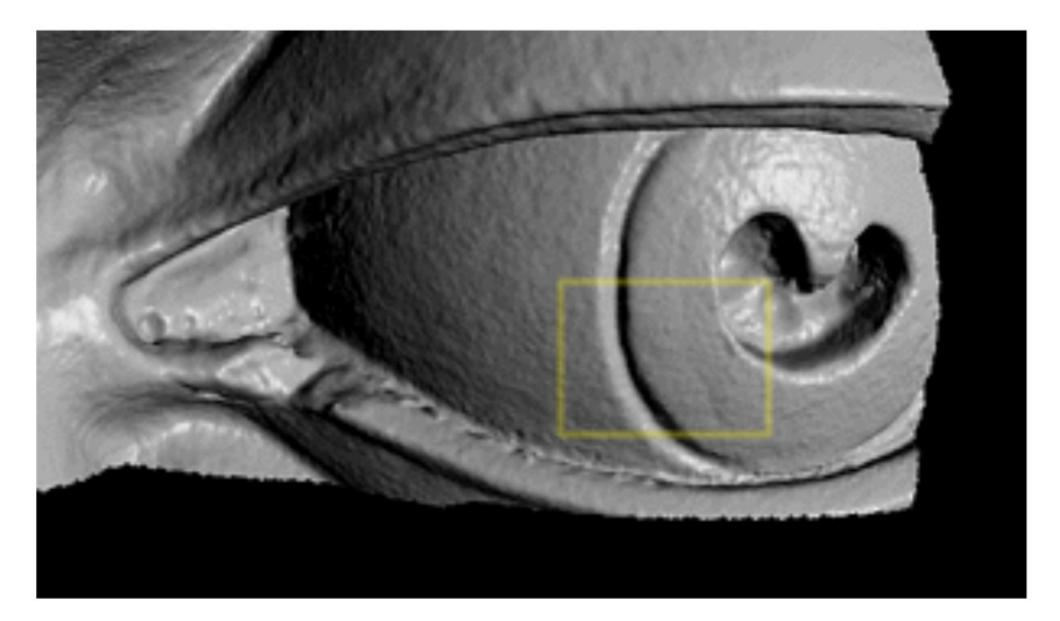
The Digital Michelangelo Project, Levoy et al.



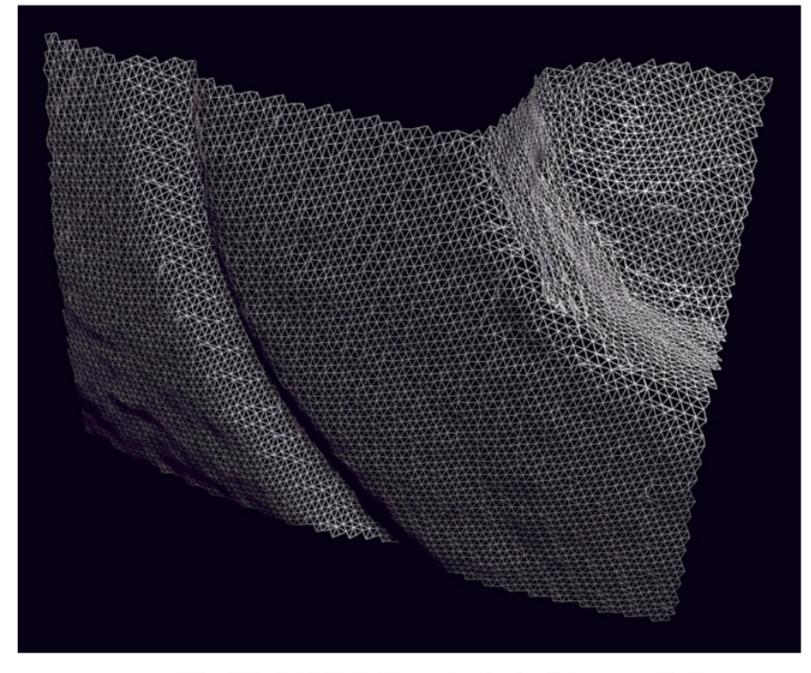
The Digital Michelangelo Project, Levoy et al.



The Digital Michelangelo Project, Levoy et al.



The Digital Michelangelo Project, Levoy et al.



The Digital Michelangelo Project, Levoy et al.

## Summary – Stuff You Need To Know

#### Camera models

- intrinsic and extrinsic parameters
- camera matrix
- camera-to-world transformation/camera-to-camera transformation

#### **Essential and Fundamental Matrices**

- How E is derived and relates to F
- How to solve for **E** or **F** using the 8-point algorithm
- How to rectify both images to be parallel
- How are the epipoles computed?

#### Stereo Matching

- How does block matching work?
- How are 3D points computed once you have a rectified stereo camera?
- Given the intrinsics, disparity, and baseline, how is a 3D point computed?

# References

Basic reading:

- Szeliski textbook, Section 8.1 (not 8.1.1-8.1.3), Chapter 11, Section 12.2.
- Hartley and Zisserman, Section 11.12.