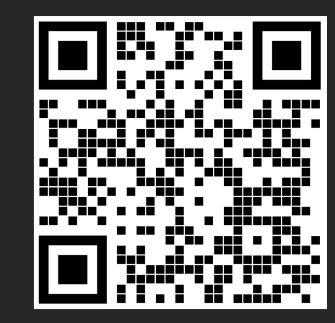
On-Sky Adaptive Secondary Interaction Matrix Calibration on the MMT

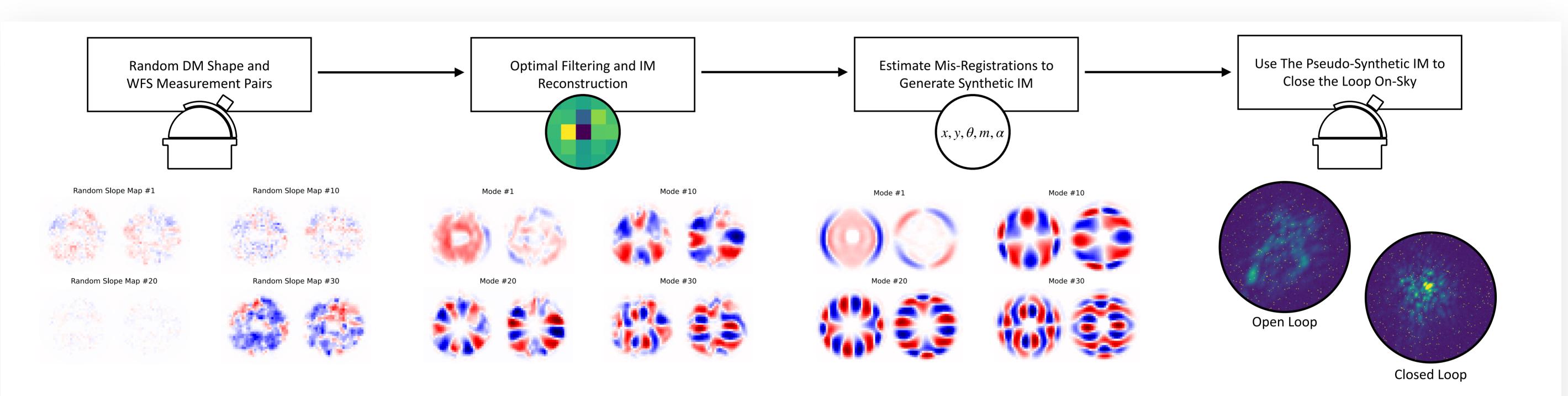
Robin Swanson ^{1, 2, @}, Jacob Taylor ^{1, 3}, Manny Montoya ⁴, Suresh Sivanandam ^{1, 3}, Amali Vaz ⁴, Dan Vargas ⁴, Grant West ⁴, Andrew Gardner ⁴, Jess Johnson ⁴, Olivier Durney ⁴, Jenny Patience ⁵, Masen Lamb ^{3, 6}, Parker Levesque ^{1, 3}

robin@cs.toronto.edu

(1) Dunlap Institute for Astronomy and Astrophysics (2) Department of Computer Science, University of Toronto (3) David A. Dunlap Department of Astronomy & Astrophysics, University of Toronto (4) Steward Observatory (5) ASU School of Earth and Space Exploration (6) Gemini Observatory, Southern Operations Center

With the commissioning of the refurbished adaptive secondary mirror (ASM) and (visible and infrared) pyramid wavefront sensors (WFS) for the 6.5-meter MMT Observatory under way, special consideration had to be made to properly calibrate the mirror response functions to generate an interaction matrix (IM). Like many upcoming extremely large aperture telescopes (ELTs), the MMT lacks a point in the optical path to place a calibration source to accurately sample the ASM's actuator response functions. We show how the DO-CRIME^[1] and SPRINT^[2] algorithms were successfully implemented at the MMT to extract an IM from on sky data and match them to a mis-registration accurate synthetic IM. We also present improvements to their base algorithms, greatly improving robustness to noise as well as errant actuators. Our ultimate goal is to provide a 100 mode pseudo-synthetic calibration with under 10 minutes of on-sky time.These methods have been validated both on an optical bench AO system as well as preliminary on-sky results from the MAPS (MMTO Adaptive optics exoPlanet characterization System) project on the MMT.





Step 1: Slope measurements created by randomly poking the ASM during on-sky

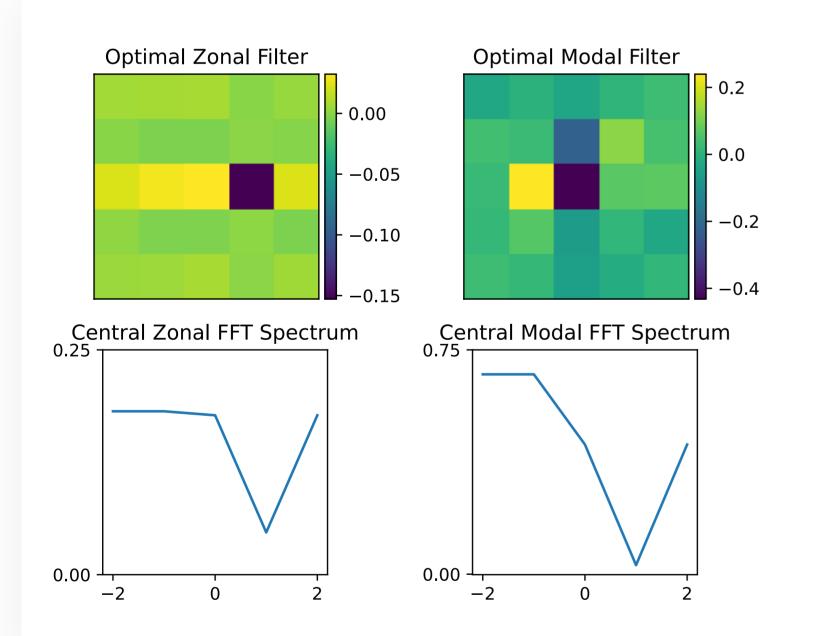
Step 2: Our optimal filter applied to the DO-CRIME method recovers an initial noisy IM **Step 3:** SPRINT aligns our noisy IM with simulation to estimate mis-registrations and

Early Results: First light results from closing the loop on MMT with our synthetic

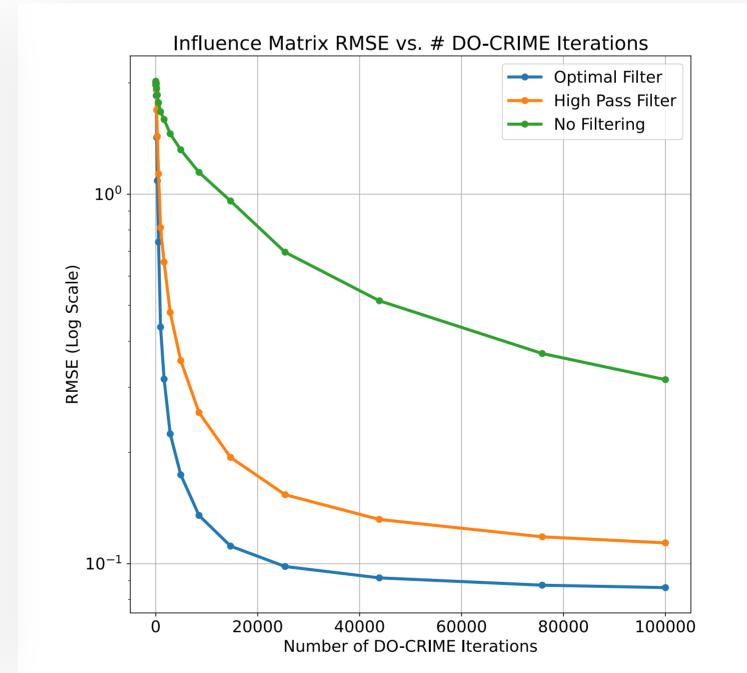
open-loop operation

create a synthetic IM

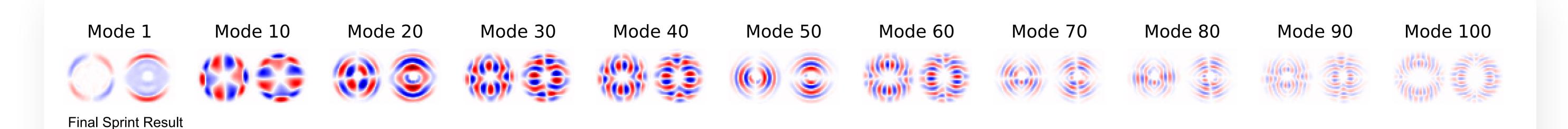
IM on MIRAC-5. More results soon to come

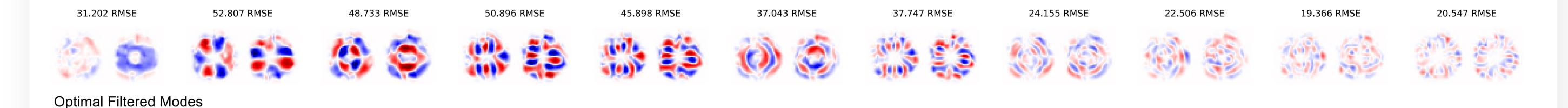


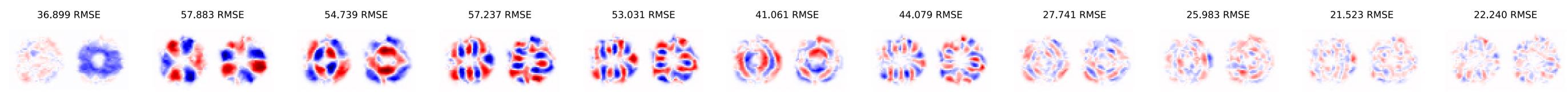
Optimal Filtering: The optimal filter for our DO-CRIME data was generated by carefully creating a "ground truth" IM on our bench and optimizing the filtering method used on our raw slope data. Dubbed the Crime Wave, our filter works equally well on optical bench and on-sky MMT data



Data efficient: Due to our improved filtering and optimization methods, we require fewer measurements of random slopes during on-sky operation to identify the mis-registrations required to generate the synthetic IM with the SPRINT method







High Pass Filtered Modes

On-sky Results: Comparison of recovered modes from raw-data to optimized SPRINT

Citations:

[1] DO-CRIME: Dynamic On-sky Covariance Random Interaction Matrix Evaluation, a novel method for calibrating adaptive optics systems — O. Lai (2020)

[2] SPRINT, system parameters recurrent invasive tracking: a fast and least-cost online calibration strategy for adaptive optics
CT. Heritier (2021)





